# On a possible explanation of Livingston-Penn effect

Yu. A. Nagovitsyn (Pulkovo Obs.) A. A. Pevtsov (NSO) W. Livingston (NSO)

## Long-term variations of magnetic field strength in sunspots



Compiled synoptic data sets (1957-2012) from 7 observatories of the former USSR (*Pevtsov etal – ApJL*, 742, 2011): **cyclic variations** 



#### Observatories of the former USSR

Crimean observatory (CRAO): continuation of USSR sunspot MF strength obs. program in 1998-2012

http://swc.crao.crimea.ua



# CRAO: "classical" observational method:

-Combination of ¼ wavelength plate and polarizer

-Measurement uncertainty – 100G

-Sunspots and large pores (polarity and field strength), small pores – polarity only

-Maximal coverage of visible sunspots



### **Example of Daily Observations**



11N = 1100 Gauss (N polarity)

## Monthly means of daily **strongest** sunspot magnetic field strengths



## LP data (with a linear correction) among yearly values of the FSU observations (strongest MF)



## Populations of sunspots (main spots in groups), Kislovodsk synoptic data



## Four types of sunspots (combinations of two populations)



SS (only small spots),  $A \le 17$  mph SL (small spots mainly+some large spots)  $17 < A \le 58$ LS (large spots mainly+some small spots)  $58 < A \le 174$ LL (large spots only)  $A \ge 174$ .

- SS: H < 800-1000 Gauss
- SL:  $H = 1000 \div 1700$
- LS:  $H = 1700 \div 2700$
- LL: H > 2700

#### Fractions of small and large sunspots



Fraction of small spots increase

### Conclusions

- No **centennial decline** in strongest magnetic field strength in the last years is observed.

- An average measured value of sunspot MFs by Livingston-Penn effect may be regulated by a current relative composition of small and large sunspots.

- Combination of different data shows that magnetic field measurements exhibit **cyclic** variations only.

- The last conclusion can be verify by the long-term Mount-Wilson observations (from 1917)