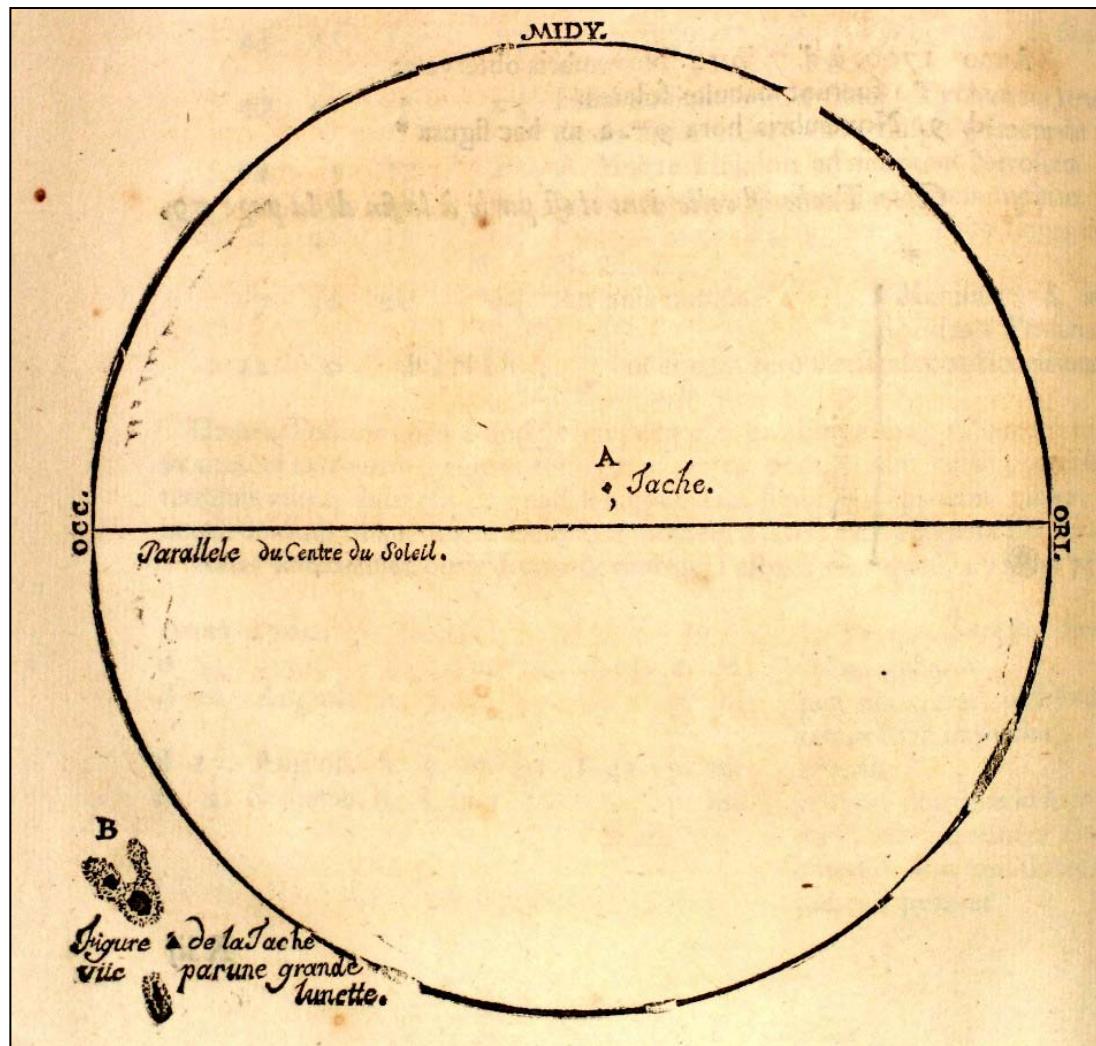


# Historical Sunspot Database

J. M. Vaquero, Universidad de Extremadura, Spain



Sunspot observed  
by Cassini and  
Maraldi from  
Montpellier (29  
March 1701)

# OUTLINE

1. How can we construct a **new version** of SSN?
2. Sunspot database **now**
3. **Some examples** of additions or corrections.

The onset of the Maunder Minimum

Drawings by Staudacher (Rainer Arlt)

Madrid SSN

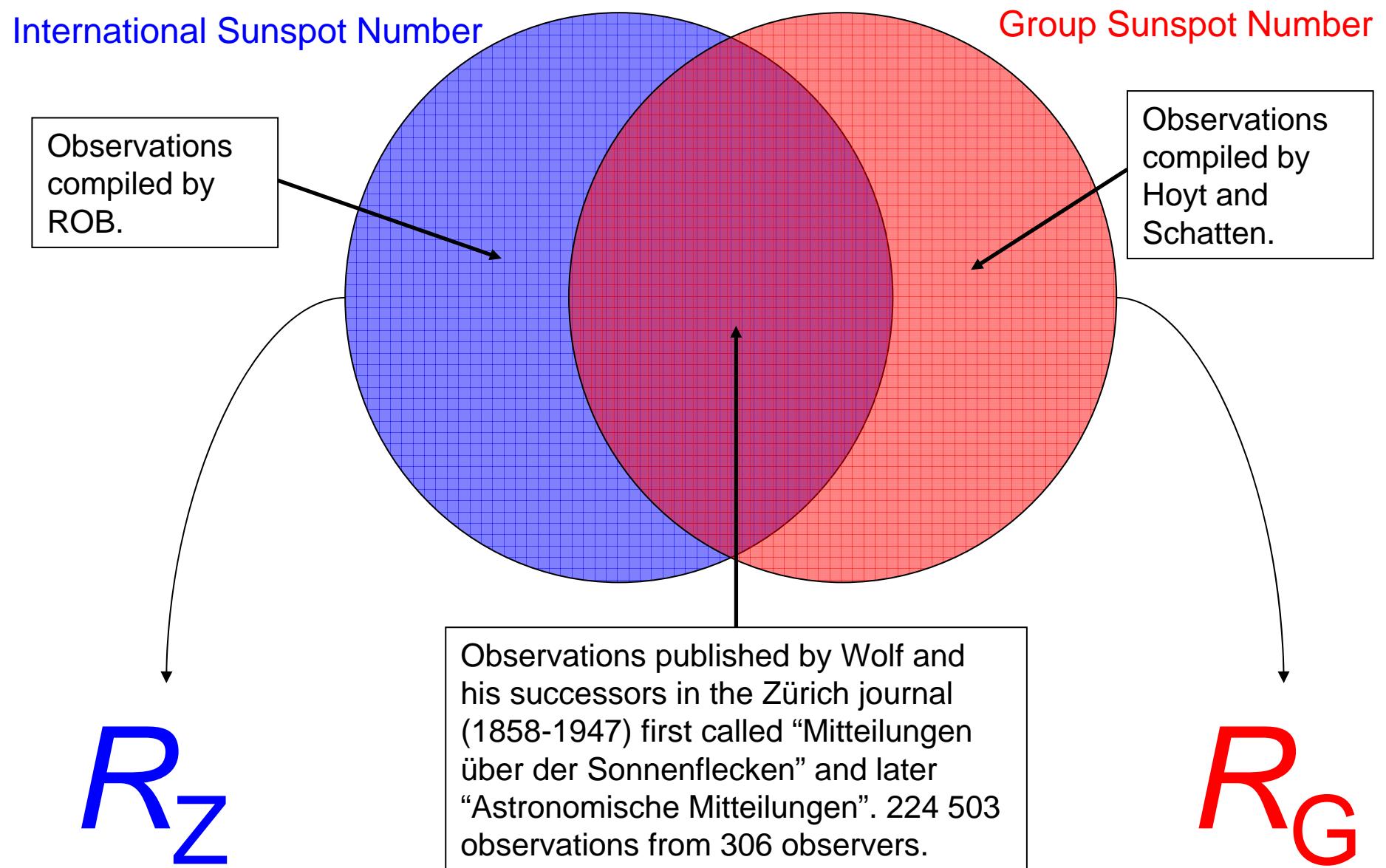
4. **Quality control**
5. Leif's **breakpoints**
6. Some **proposals**

Addenda: **Annual Averages** of SSN!!!

# 1. How can we construct a new version of SSN?

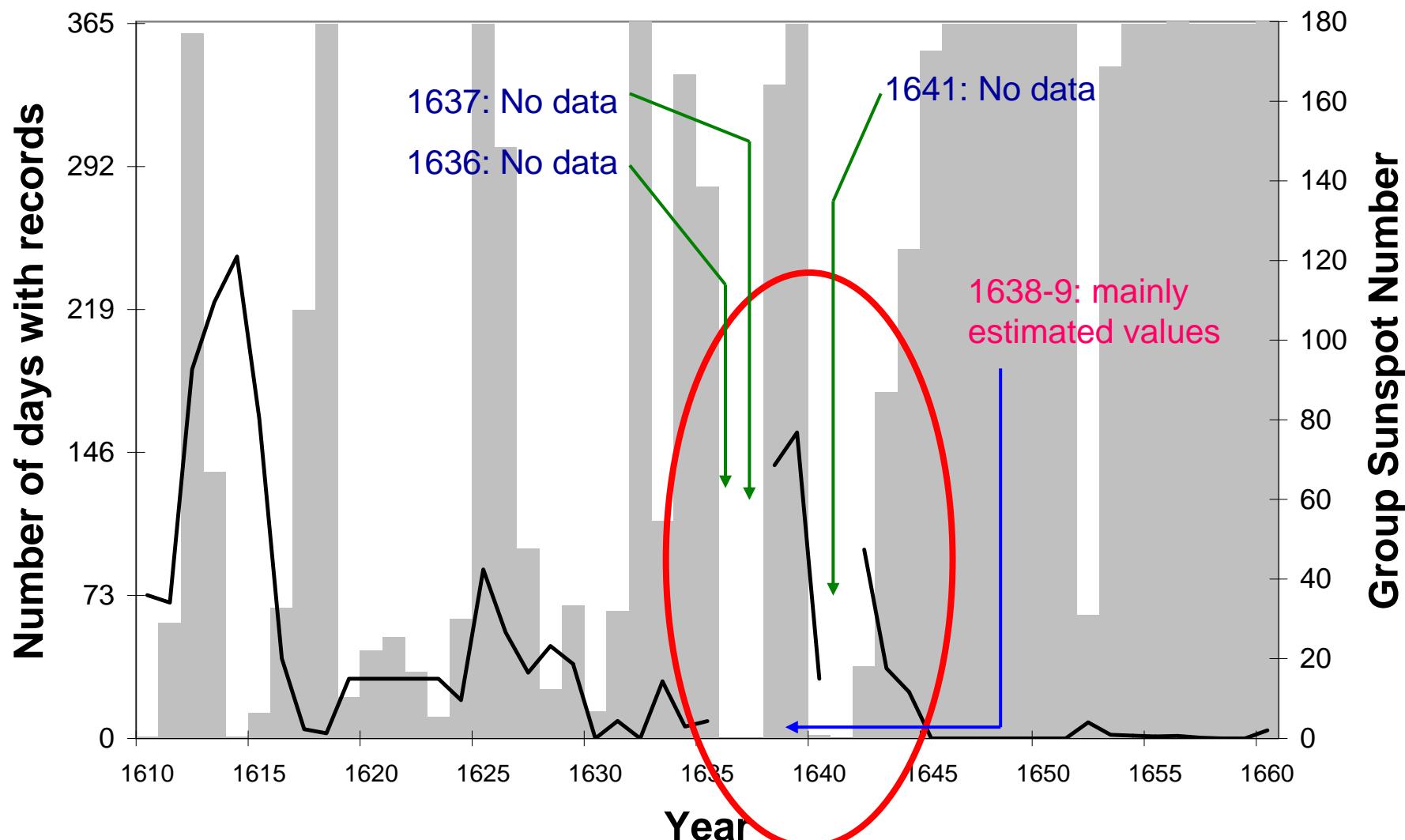
- I. Improve the database with old and new observations
  - I. Pending tasks (Hoyt & Schatten, 1998)
  - II. Rediscovering “lost” records
- II. Develop a new methodology to reduce all observations
  - I. Computing classical calibration constants ( $k$ )
  - II. New statistical techniques
- III. Check our series using different tools
  - I. Geomagnetic records (c. 1835) and other proxies
  - II. Statistical techniques
- IV. Publish and maintenance of data

## 2. Sunspot database now



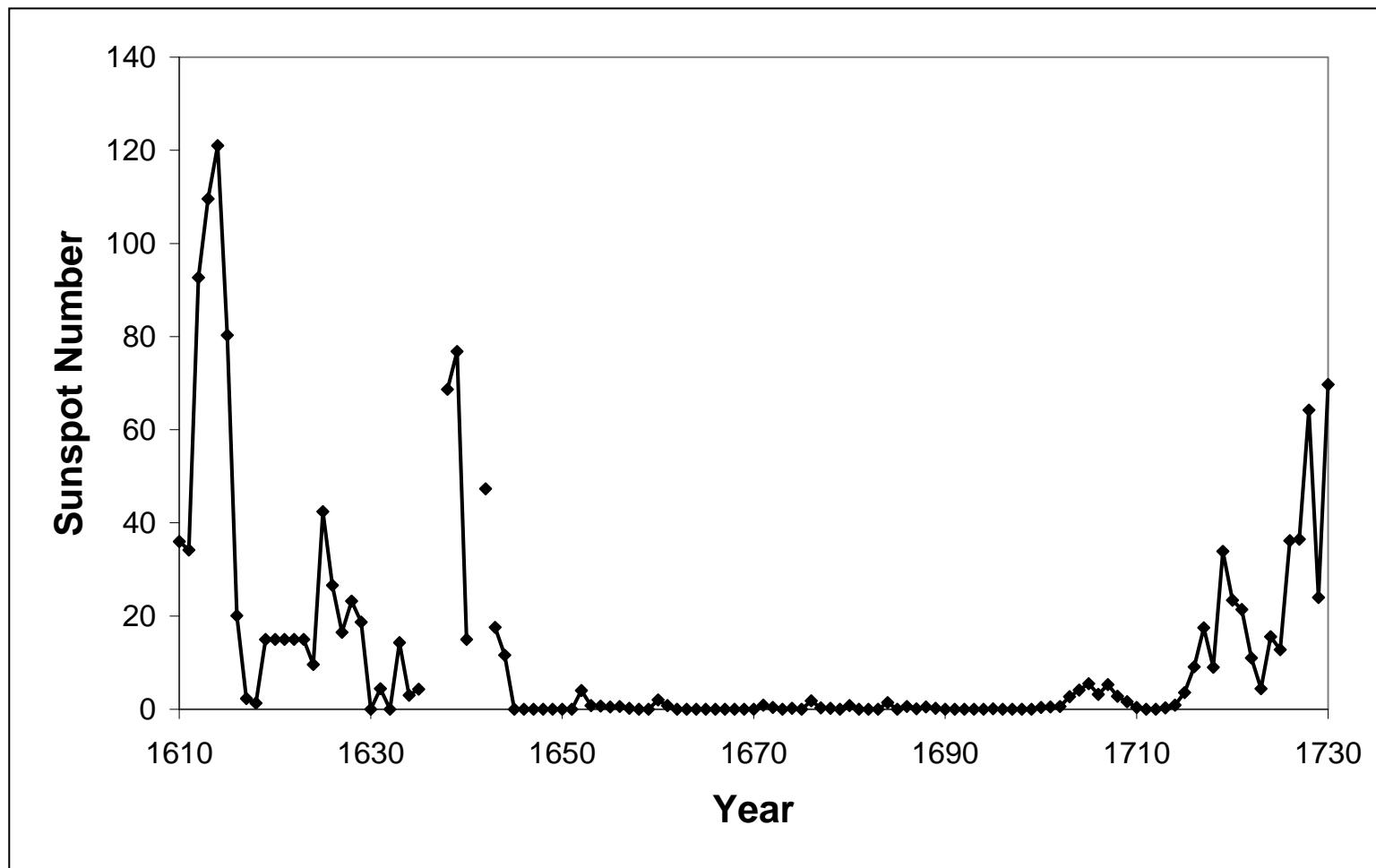
### 3. Some examples of addition or corrections

#### 3.1. The onset of the Maunder Minimum



Accepted general scenario for Maunder Minimum (Usoskin, 2008):

- (1) transition from the normal activity to the deep minimum was sudden,
- (2) a 22-year cycle was dominant in sunspot, and
- (3) the recovery of the sunspot activity from the deep minimum to normal activity was gradual.



(1) We have added the **Marcgraf sunspot records**



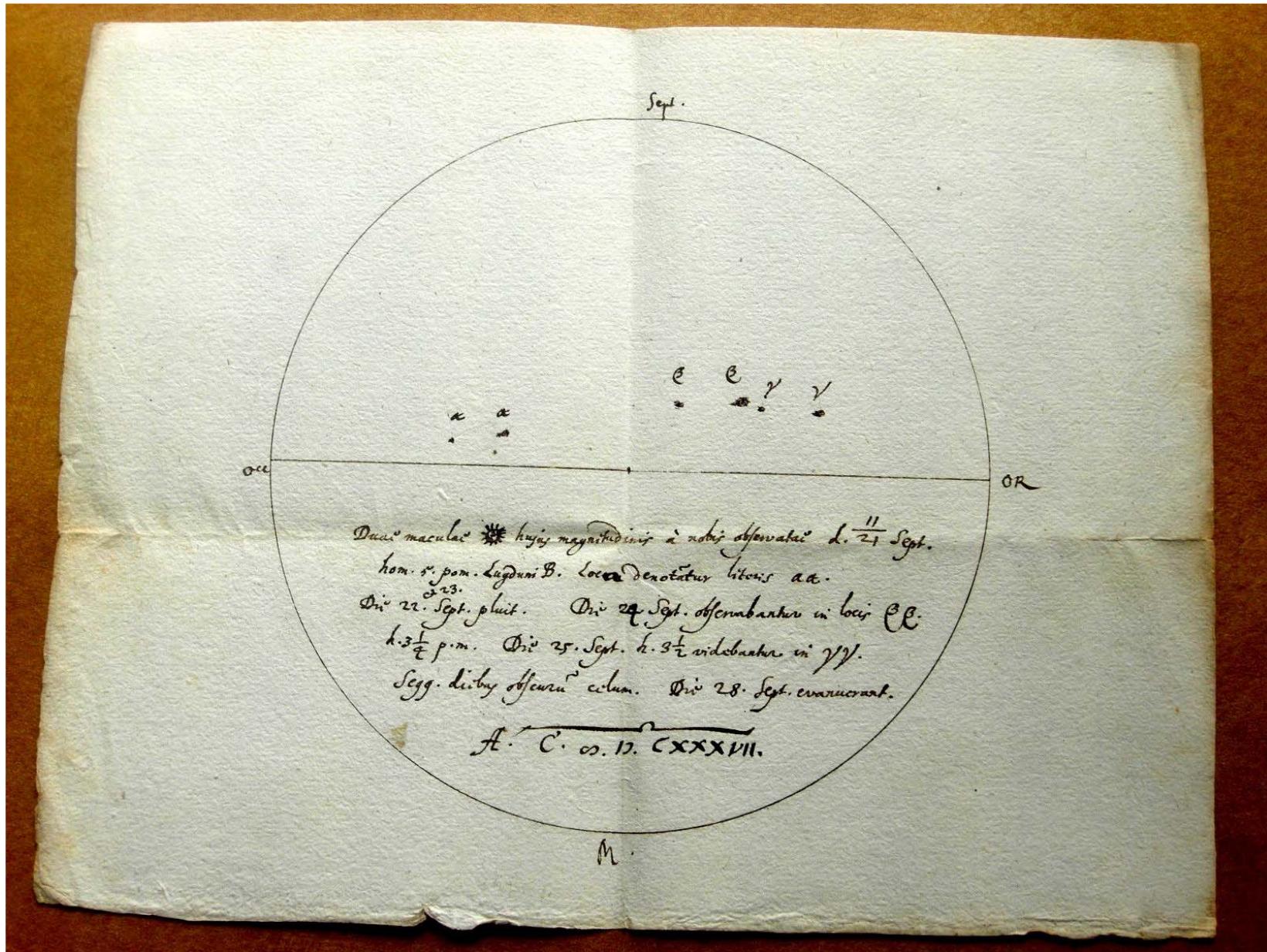
Die 1. Feb. abwart.	
bis 2. Feb. Alt. O. Med. 21-8. in Kasten 21-136 versteckt.	
Die 3. o. 4. Feb. abwart.	
Die 5. Feb.	
Alt. O. Med. 22-6. Rauh 3 violett in Kasten 21-136 versteckt.	
In ab. Pergament 25-18-02 ab 25-20.	
Die 6. Feb.	
Alt. O. Med. 22-6. Rauh 3 violett in Kasten 21-136 versteckt.	
Die 7. Feb.	
Alt. O. Med. 24-40.	
Die 8. Feb.	
Alt. O. Med. 24-40.	
Die 9. Feb. abwart.	
In ab. jefer ay II - Or. 51-2-20 <small>(ab 21-136)</small>	
Kapf in Markstab Perg. 14-37 ab 21-136 versteckt.	
Die 10. Feb.	
Alt. O. Med. 23-40.	
Die 11. Feb.	24-0.
Alt. O. Med.	
Die 12. Feb. abwart.	
in Kasten 21-136 versteckt.	
Alt. O. Med. 30-18.	
ab 25-12 in Or. (ab 21-136) 8.	
ab 25-12 ab 25-13-50.	
in Markstab Perg. 22-34	
Die 13. Feb.	
Alt. O. Med. 24-40.	
Die 14. Feb. abwart.	
In ab. jefer ay II - Or. 53-19.	
ab 25-19.	
ab 25-19 ab 25-50.	
ab 25-19 ab 25-37.	
ab 25-19 ab 25-4.	
Die 15. Februar.	
Die 16. Februar.	
Alt. O. Med. 25-41.	
Verg. in ab. jefer ay II - Or. 49-36-20.	
ab 25-20 ab 25-22.	
ab 25-20 ab 25-4.	
ab 25-20 ab 25-7.	
Die 17. Februar.	
Die 18. Februar.	
Alt. O. Med. 25-41.	
ab 25-20 ab 25-22.	
ab 25-20 ab 25-56.	
ab 25-20 ab 25-56.	
ab 25-20 ab 25-9.	
ab 25-20 ab 25-14.	
No. 1. A. in Kasten (ab 25-20 ab 25-22)	
ab 25-20 ab 25-22.	
ab 25-20 ab 25-22.	

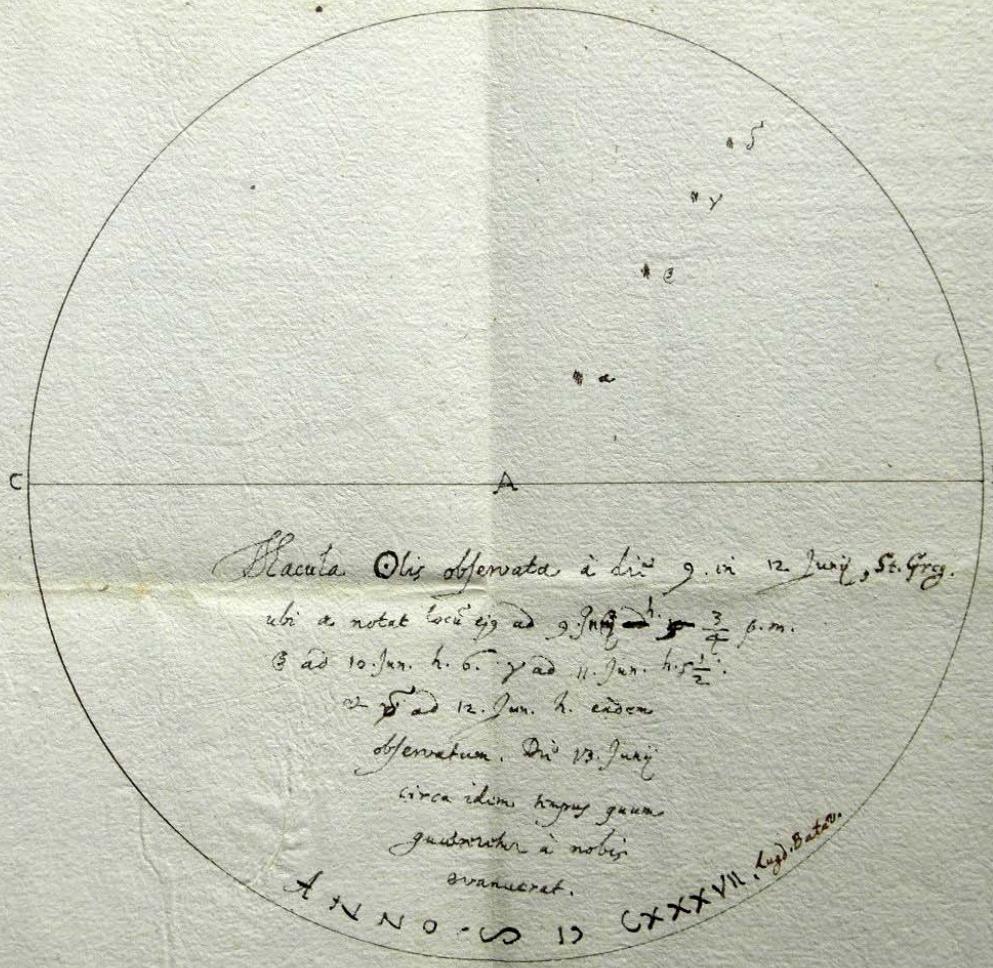
Die 3. & 4. Feb. wurden  
Die 5. Febr.  
Alt. O. Medianus 22. 6%. Maxilla 3 mit  
in Sols hoc m. Vagi h.  $\frac{1}{2}$ . v  
In. alt. Prognathia or. phya 25. 18. c  
25. 27. 20. Sols 10. 17.

a b

Macula ☽ à nobis observata M. octob. A.C. 13 CXXXVII. Lugduni Bat.  
Die 10. Octob. st. hinc. nalla inveniatur macula. Die 11. erat observata cœli. Die  
12. Octob. h. 2 $\frac{1}{2}$  pom. macula primo inveniatur in a. Die 13. Octob. h.  
3. pom. visa nobis in b. ad exigua, et minor quam pridie. Die 14. Octob.  
hinc sole. Die 15. h. 3 $\frac{1}{2}$  pom. evanescat, nec inveniatur ultra.

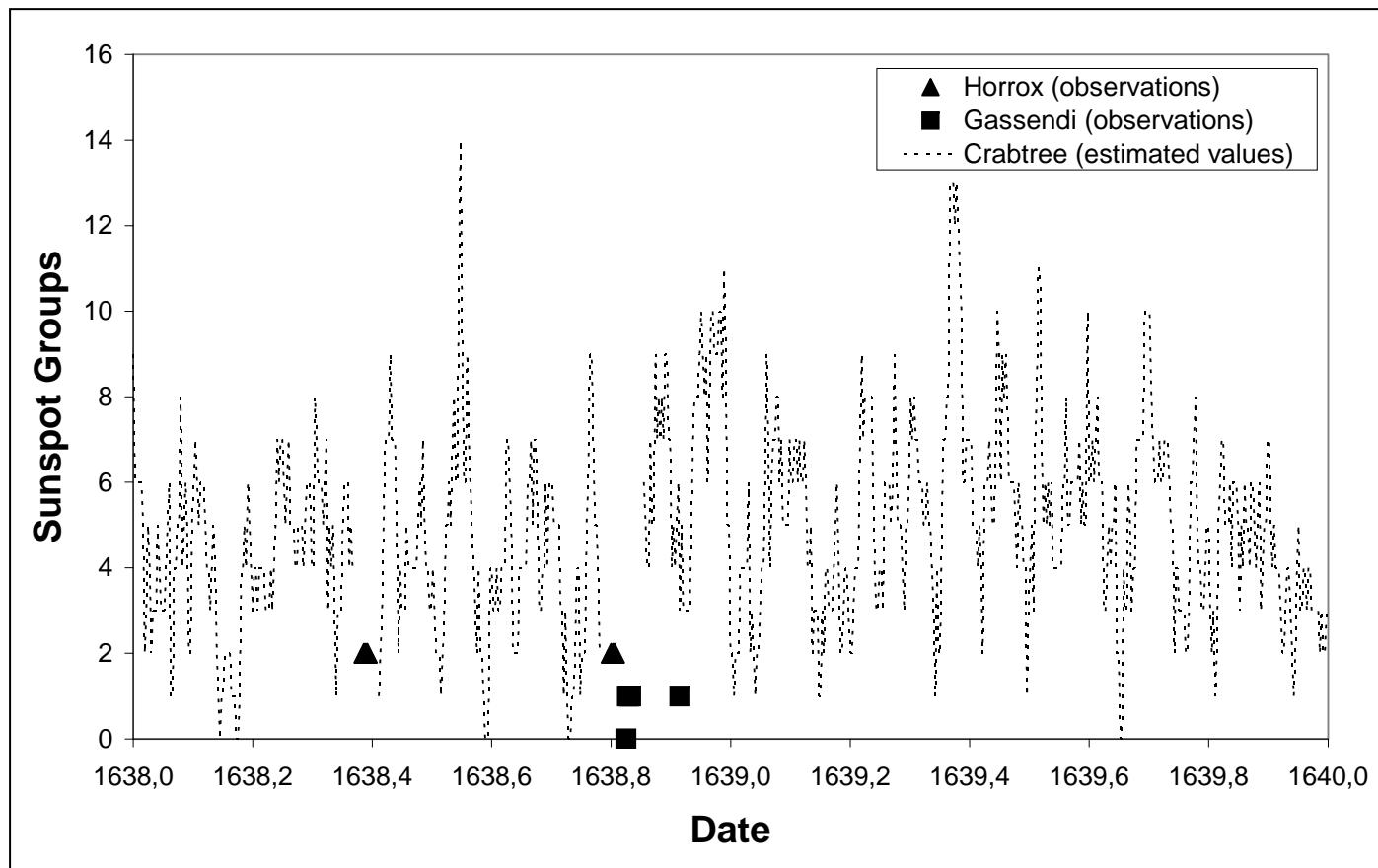
o. 8. 43  
m. 11. m. 11. 11.  
G. 11. 11. N





(2) We have **eliminated the estimated** (not observed) values from Crabtree's comments (1638-1639).

Hoyt and Schatten (1998) wrote in their Bibliography: "According to a letter by Crabtree the average number of spot groups seen in 1638 and 1639 were 4–5 per day. The database has Greenwich fill values to give 4–5 groups per day. This substitution technique was used to simplify the analysis. This is the only place in the entire database where we do this type of substitution".



(3) We have corrected the dates and the numbers of sunspot groups of Horrox observations in HS98 (**from Julian calendar to Gregorian Calendar**).

JEREMIÆ HORROCCII,  
LIVERPOLIENSIS ANGLI, ex Palatinatu  
LANCASTRIÆ,  
OPERA POSTHUMA;  
viz.  
Astronomia Kepleriana, defensa & promota,  
Excerpta ex Epistolis ad *Crabtraum* suum.  
Observationum Cœlestium Catalogus.  
Lunæ Theoria nova.

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Accedunt  
GUILIELMI CRABTRÆI, Mancestriensis,  
Observationes Cœlestes.

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In calce adjiciuntur  
JOHANNIS FLAMSTEDII, Derbyensis,  
De Temporis Aequatione Diatriba.  
Numeri ad Lunæ Theoriam Horroccianam.

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LONDINI,  
Typis GUILIELMI GODFREDI, Impensis J. MARTYN Regalis  
Societatis Typographi, ad insigne Campanæ in Coemeterio  
D. Pauli, Anno Domini M. D. C. LXXXIII.

*Et Observationes cœlestes.* 309  
putaveris te alterum ad Septentrionem, alterum ad Austrum consper-  
xisse.  
Maii 22, h. 3. P. Radios Solares intromisi per Tubum Opticum  
in cameram obscuratam, (densis tenebris non opus est:) atque in lim-  
bo orientali, prope viam regiam (ut puto,) vidi duas nigras maculas.  
Major distabat à Solis margine min. 3'. Maculae diameter longior e-  
rat quasi min. 30", brevior quasi min. 20", eratque in forma Ovali.  
Maii 23, h. 3. P. Major duarum absuit à margine min. 1, 20".  
Maii 24, h. 3. P. absuit min. 0, 25"; jāmque facta erat minor  
quām reliqua, quæ à margine distabat min. 1'.  
In magna distantia erant hæ maculae colore cœruleo, cum rubro mi-  
sto. Et circa Solis discum erat circulus cœruleus, inter exteriorem ru-  
brum, & viridem interiorem, sicut in Iride; sed pro varia vitri positio-  
ne subinde variabantur hi colores.  
Eset macula illa, si rotunda, Venere major, possitque Cometae in-  
star esse. Si à Sole projiceretur quantum inde Terra distat, ad instar  
stellæ appareret, nisi forsan à Terrâ nimis distaret.  
Post illos tres dies maculas nullas vidi. Tempus est ut reverterentur  
cædem, sed nubes impediunt observationem.  
Est autem Tubus hic meus ex vulgaribus unus, pretii 2*l.* 6*d.*, con-  
tulit tamen cum duobus aut tribus aliis, quos mutuo habui, sed meo  
(quantum ego judico) inferioribus.  
Hunc modum existimo egregium fore ad observandas Eclipses. Ad-  
mittit enim discum Solarem tantæ magnitudinis in parva distantia, ut  
ferè minuta secunda possis observare; atque lucem ab umbrâ accuratè  
distinguit, si ad justam longitudinem educatur.  
Stellas fixas dum contueor, nihil video aliud quām radios undecunq;  
emissos, pro vario vitri positione situs mutantates.  
Mars videtur ejusdem quasi magnitudinis cum Jove: *Keplerus* ta-  
men & *Lansbergius* multo majorem faciunt.  
Si Mars sit Terrâ major, oportet Solis parallaxin multò minorem  
esse quām vult *Keplerus*.  
  
*Ex Epist. Julii 25. 1638, Toxtethæ:*  
Postquam te viderim, nihil à te accepi, necdum ad te scripsi quic-  
quam.  
*Lingomontanum* tandem nactus sum. Habet ille multas observatio-  
nes Planetarum omnium, præsertim in oppositione Solis, sed breviter  
descripas.

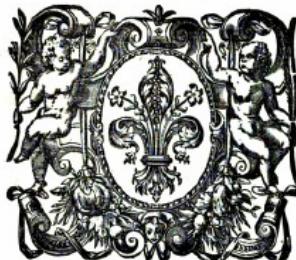
(4) We have eliminated one spurious observation by Gassendi on 1 Dec 1638.

PETRI  
**GASSENDI**  
 DINIENSIS  
 ECCLESIAE PRÆPOSITI.  
 ET IN ACADEMIA PARISIENSI  
 MATHESEOS  
 REGII PROFESSORIS  
**ASTRONOMICA,**

V I D E L I C E T

- I. Institutio Astronomica cum Oratione Inaugurali.
- II. Observations Cælestes.
- III. Mercurii in Sole visus & Venus inuisa.
- IV. Noum Stellæ circa Iouem visæ.
- V. Solstitialis altitudo Massiliensis.

TOMVS QVARTVS.  
*CVM INDICIBVS NECESSARIIS.*



LVGDVN,  
 Sumptibus LAVRENTII ANISSON.  
 & IOANNIS BAPTISTÆ DEVENET.

M. DC. LVIII.  
*CVM PRIVILEGIO REGIS.*

Commentary.

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¶ circiter medium visputa, & Procyon 3500—800. hoc est  
 ♀ & Arcturus 3500—41200. hoc est 17. grad.1. min.  
 D 18 17. Manè, cum die superiore ab antelucano visque tempore in ipsam visque no-  
 men vixisset, superiore deinde nobis, que Famulo hinc ad S. Vincentium progreſſo vi-  
 sis à 4. marzeti quibusdam rubrum hianibus : & Aggrato tamen Calo Aquis Sextis fecitno,  
 distinatum inter le

2& Spica mp 3700—41982. hoc est 31. grad.1. min. idque hora 6.

2& Spica mp 3700—41714. hoc est 16. grad.1. min.

2& & Arcturus 3700—41530. hoc est 40 grad. min. 15 $\frac{1}{2}$ .

D 18 29. Manè, nihil visum Aquis-Sextis ( vi neque etiam die superiore tam Aquis-  
 Sextis, quam Dinitis ) Famulo autem ad Sandrum Vincientum, distinatur Cœlo lati-  
 pulcro

♀ & Spica mp 1000—41310. hoc est 28. grad. min. 47 $\frac{1}{2}$ . Alta ♀ 6. grad.5. min.  
 Arcturus & Spica mp 500—41240. hoc est 32. grad. min. 53 $\frac{1}{2}$ .

2& ♀ 1500—2960. hoc est 1. grad.4.8. min.

2& 2& 3200—3010. hoc est 0. grad. min. 55 $\frac{1}{2}$ . forte—3300. hoc est 1. grad. min. 1 $\frac{1}{2}$ . Alto 34  
 5. grad.40. min.

(opinor non 2&) & ♀ 3500—1740. hoc est 3. grad. min. 51 $\frac{1}{2}$ .

& 2& iterum 3100—3110. hoc est 0. gr. m. 51 $\frac{1}{2}$ . Reruit hanc distinatum optimè acceptam.

D 18 30. Manè, Aggrato, Aquis-Sextis, distinuntur inter se

& & Austrina lnx 3900—1514. hoc est 14. grad. min. 3 8 $\frac{1}{2}$ . Audita iam fuerat hora 6.

& lnx eadem 2& 3900—1211. hoc est 8. grad.47. min.

& Spica mp 3900—41271. hoc est 14. grad.6. min.

& lnx Borea 2& 3900—1790. hoc est 11. grad. min. 51 $\frac{1}{2}$ .

& ♀ 1000—1197. hoc est 1. grad.15. min.

& Spica mp 3900—41569. hoc est 30. grad.13. min.

& ♀ 1000—1690. hoc est 1. grad. 43. min.

Eodem Manè, Famulo ad S. Vincentium, distinuntur inter se

Lncs 2& 3100—1770. forte pro 1707. hoc est 9. grad.11. min.

Spica mp & lnx Austrina 1000—41770. forte pro 1707. hoc est 11. grad.11. min.

Spica mp & Arcturus 3500—41230. hoc est 12. grad.56. min.

2& & Spica mp 3500—41900. hoc est 10. grad.8. min. Altitudi Arcturi 45. grad.3. min.

2& & ♀ 3100—5104. hoc est 1. grad. min. 2 $\frac{1}{2}$ .

& Spica mp 3500—41280. (opinor pro 280.) hoc est 34. grad.8. min.

& ♀ 1500—2780. hoc est 1. grad. min. 45 $\frac{1}{2}$ .

& 2& 3500—3101. hoc est 1. grad. . min. Alta 2& 8. grad.0. min.

& ♀ iterum 3100—2760. hoc est 3. grad. min. 1 $\frac{1}{2}$ .

**M**ENSE DECembR, Die 1. Manè (fuerat Die 1. Cœlum strobique obductum)

distinut Aggrato Aquis-Sextis

2& ♀ 3900—3400. hoc est 2. grad. min. 3 1 $\frac{1}{2}$ . Sic nec?

& lnx Austrina 3900—2910. hoc est 15. grad. min. 38 $\frac{1}{2}$ .

& lnx Borea ipsi perpendicularis 1900—1161. hoc est 13. grad. min. 31 $\frac{1}{2}$ .

& 2& 1900—2980. hoc est 4. grad. min. 43 $\frac{1}{2}$ .

& ♀ 3900—3049. hoc est 4. grad.21. min.

& Spica mp 3900—41000. hoc est 32. grad. 57. min.

& Spica mp 3900—41960. hoc est 32. grad. 24. min.

Eodem Manè Famulo ad S. Vincentium, distinuntur inter se

Lncs 3100—1834. hoc est 8. grad. 54. min. Debuti 9. grad. min. 8 $\frac{1}{2}$ . An alia Stella.

Spica mp & Arcturus 3500—2420. hoc est 32. grad. min. 51 $\frac{1}{2}$ .

2& & ♀ 3500—3410. hoc est 0. grad. min. 17 $\frac{1}{2}$ . Altitude Auctuti 49. grad.0. min.

2& & Spica mp 3500—41310. hoc est 32. grad. min. 34 $\frac{1}{2}$ . Alto 2& 6. grad.0. min.

2& & 2& 3500—2790. hoc est 3. grad.43. min.

& ♀ 3500—2260. hoc est 3. grad. min. 17 $\frac{1}{2}$ .

2& & ♀ iterum 3500—3410. hoc est 0. grad. min. 17 $\frac{1}{2}$ .

D 18 4. Manè (fuerat Cœlum superior obductum) distinut Aggrato Aquis-Sextis Co-

ro vigente, & frigore præcuso

& Austrina lnx 2& 3900—4120. hoc est 18. grad. 43. min.

& Borea lnx 3900—832. hoc est 16. grad.11. min.

& margo Borens, vicinalique 3900—3148. hoc est 3. grad.1. min.

& ♀ 1900—2959. hoc est 4. grad. min. 49 $\frac{1}{2}$ .

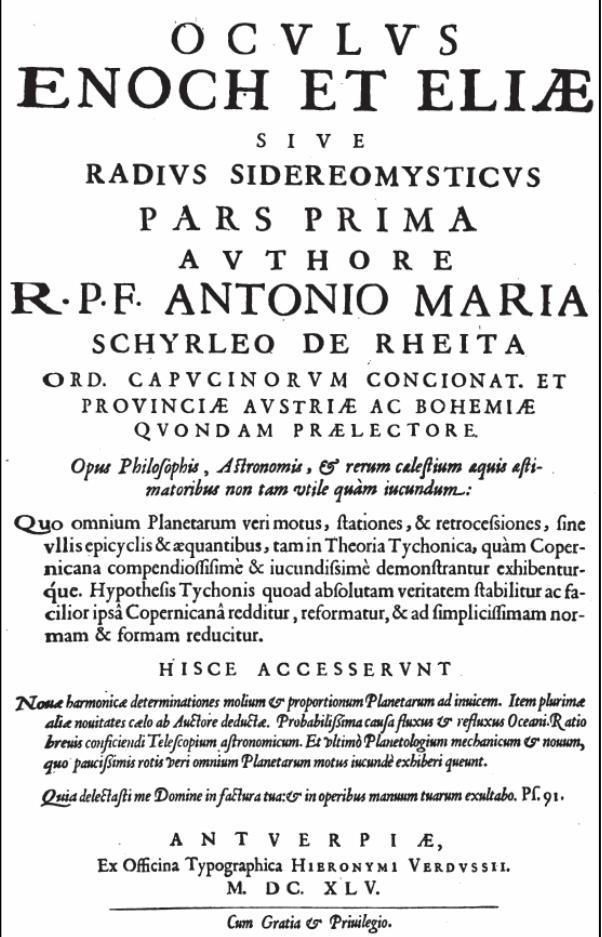
2& & Spica mp 3900—41242. hoc est 15. grad. min. 56 $\frac{1}{2}$ .

2& & ♀ 3900—3432. hoc est 1. grad. min. 21 $\frac{1}{2}$ . Meo manè hoc non vacavit. Et de his 2& ap-  
 partinet tantum.

Gassendi Observationes.

DD d D 18

## (5) We have changed the record by Rheita [1642].



### L I B E R Q V A R T V S .

242 & Sol morbo illo palloris, & seculum tristiori caligine laborarint. Idem sub Constantino Principe, & Irene contigisse ferunt circa annum Christi 797. Quorum tempore per 17, integros dies, adeò nufquam visus est Sol, adeò tenuerat radios telluri immisit, vt mundo aut omnino abreptus, aut certè radijs & gratissimâ luce spoliatus exutusque crederetur. Enimvero in valtissimo tunc oceano Oberantes, neque cursum suum per te- nebras dirigere, neque telluri insufflentes ceptum iter & negotia- tiones humanas prosequi potuerunt.

Iterum Anno 1547. per totam valtissimam Europæ plagam, Solis radij sanguineo colore adeò detleti videbantur trium dierum spatio; vt mundo ultimum iam quasi inciperent prænuntiare diem. Denique tempore, quod Rudolphus II. Augustus ex humanis abreptus fuerat, Solen per plures dies, & lumen notabilissime hebetari debilitari- que.

Quidam Solem finitem moni- fuisse putant. Horum igitur solarium prodigiis merito causam indagare licet. Aliqui putant Solem in fastu alterius montis Aethnae, aut Vesuvij incrementa sua in extimam superficiem proflare, & veluti pluviâ fauillarum inde adeò conpergi- & vndique circumdari, vt mundo inde quasi eripiatur dies, splendore omni Solis intercepto, donec eru-

Certè quod iam diximus, pro- pria experientia Colonia: Anno 1642. experti sumus: dum ingen- ti stellarum solarium turram maiorum & minorum per 14. dies & ultra sibi inuicem continuâ se- rie succendentium cum stupore, so- larum discum adeò occupare vidi- mis, vt lux eius, maxime media, & intensissima, haud leviter illis fuerit hebetata. Nam cubo opti- mo, in medio solaris disci globum perfectissimè rotundum, subni- grum,

### C A P V T S E C V N D V M .

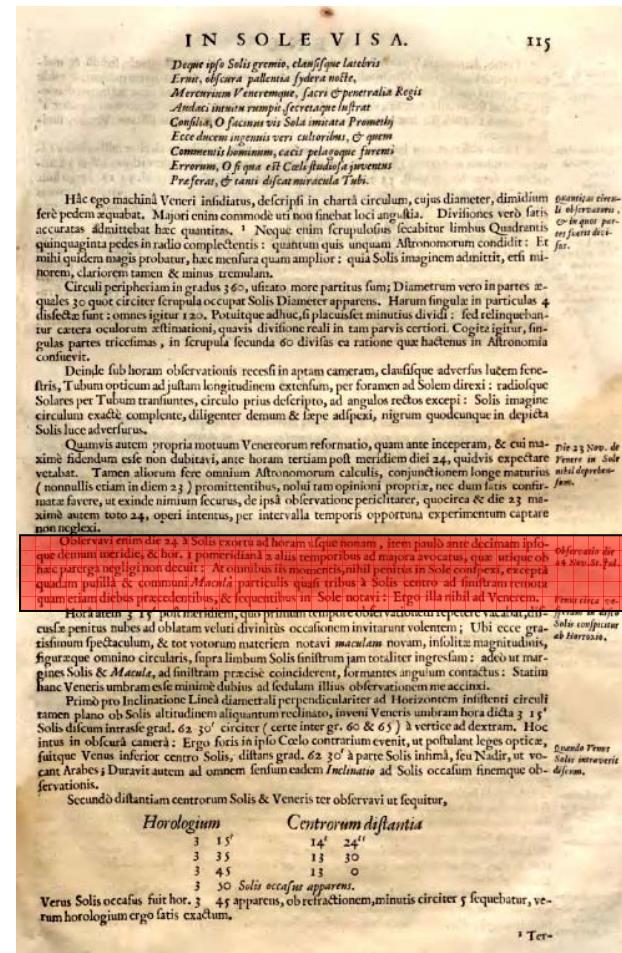
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subnigrum, pugni magnitudinem quasi excedentem conspicimus, idque directissimo aspectu; qui & per octiduum Solis haud exiguum, portione eclipsis fuit: maximali que aeri turbationes, & vrpote ventos, imbræ, & frigora in medio lunij artulit: prout crebris observationibus iam à multis annis compertum habemus: scilicet ferè semper aëris insigniores, & magis notabiles mutationes ex dictarum stellarum solarem discum subeuntum agmine contingere & entire.

Espresso perfalsum est, ma- culas illas penitiori obtutu direc- te per optimum & longiore tubum astronomicum, ( qui totum simul solarem discum discooperiat exhibeatque) conspectas, aliam quam circula- risimam & rotundam figuram ostendere, vt frequenter experti sumus. Itaque toties semper solares eclipses contingere necesse est, quoties stellæ dictæ Solem sub- eunt; subeunt autem frequentissime; ergo multò frequentiores & plures contingunt nobis solares eclipses, quam vulgus arbitratur. Sed quis oblectot alium ecliplium arcenos respectu telluris nostræ effectus haec tenus penetravit? vt quid ergo paupelli illi deceptores Astrologi, ex astris de futuris contingentibus diuinare non erubef- cunt, cum multa praestentia in astris ignorantia sidera & ita cæsis & falsis suis prognosticis procedant, Pars I.

H 2 tum

## (6) We have incorporated a sunspot record by Horrox in 4 December 1639.



non neglexi.

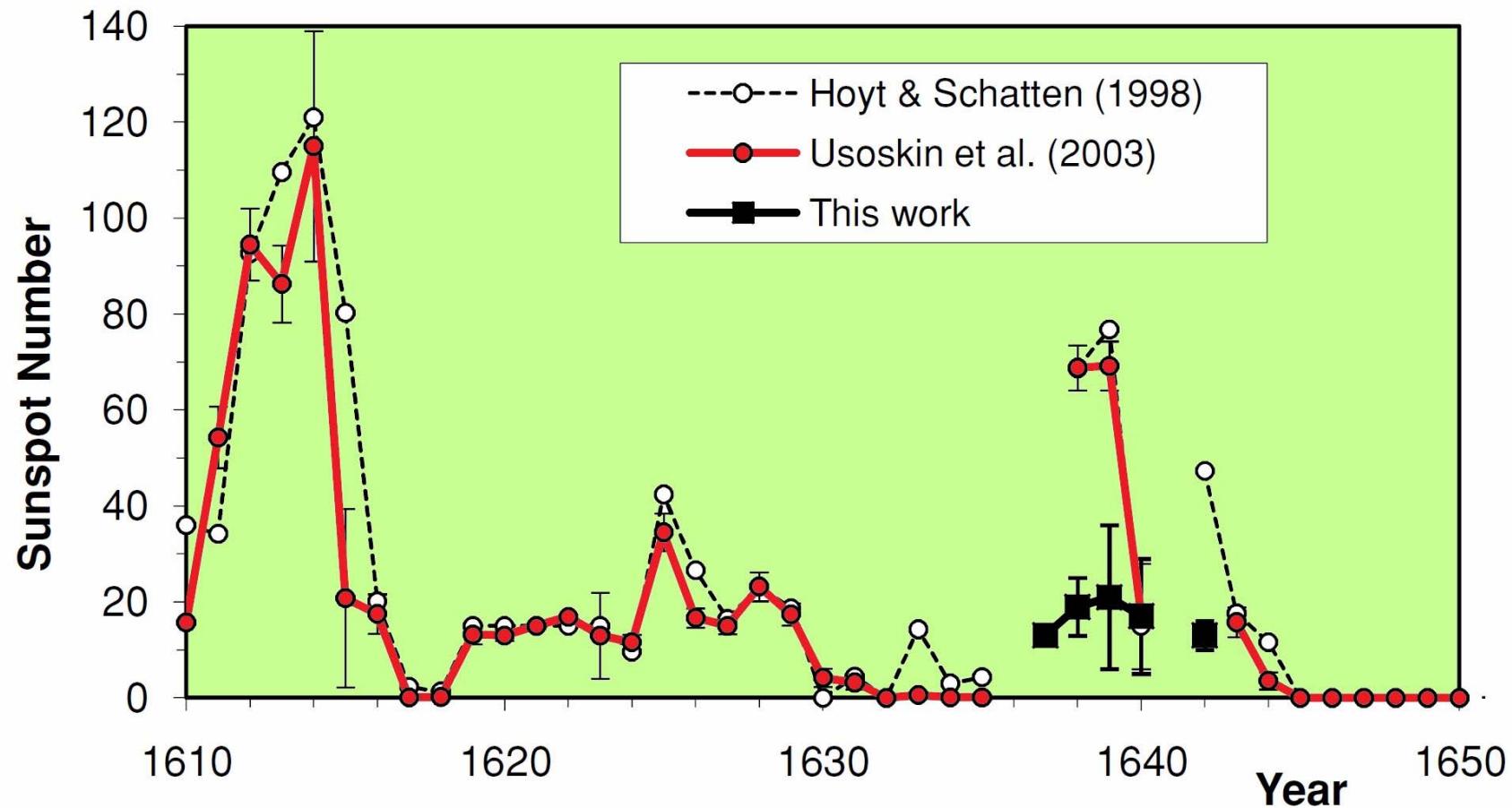
Observavi enim die 24 à Solis exortu ad horam usque nonam, item paulo ante decimam ipso-  
que demum meridie, & hor. 1 pomeridianâ 2 aliis temporibus ad majora avocatus, quæ utique ob  
hæc parerga negligi non decuit: At omnibus iis momentis, nihil penitus in Sole conspexi, excepta  
quadam pusilla & communis *Macula* particulis quasi tribus à Solis centro ad sinistram remota  
quam etiam diebus præcedentibus, & sequentibus in Sole notavi: Ergo illa nihil ad Venerem.

Horæ atem 2 15' post meridiem, quo primum tempore observationem renetere vacabat. dif-  
feram in disco

Observatio die  
24 Nov. St. Jul.

Venus circa ve-  
nerem in disco

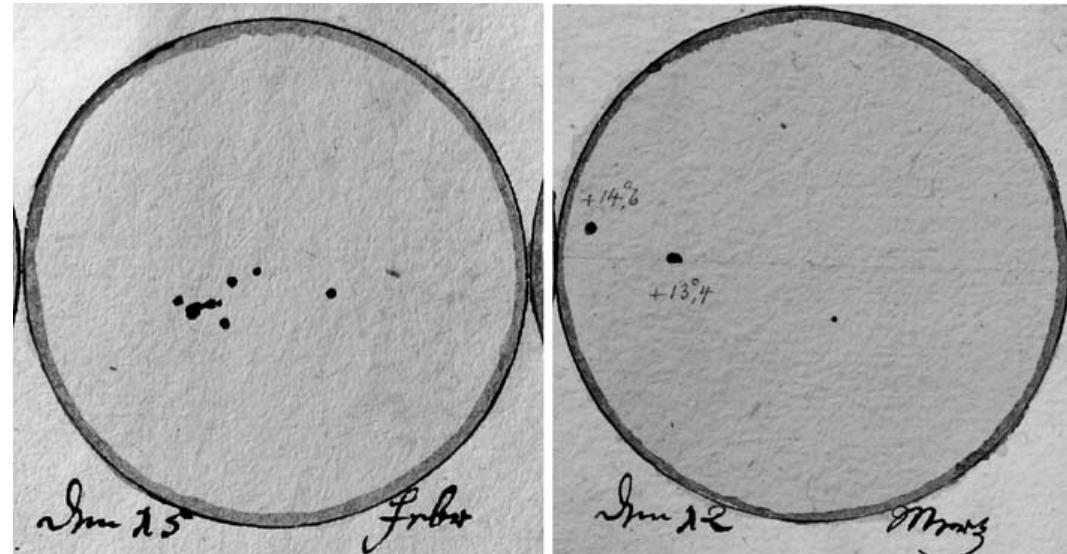
We can use a statistical procedure (Usoskin, Mursula & Kovaltsov, 2003) to reconstruct yearly group sunspot number from sparse daily observation.



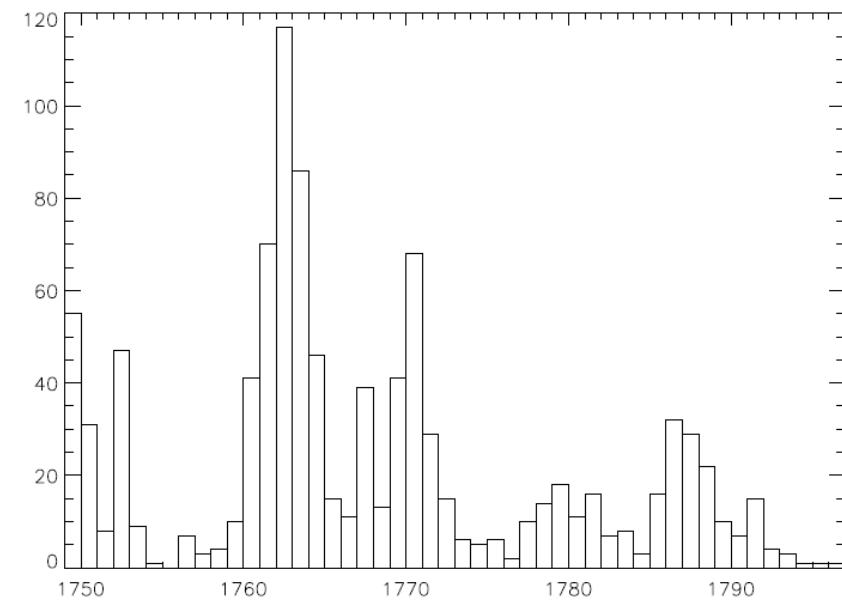
Vaquero et al. (2011), ApJ

### 3.2. Drawings by Staudacher (Rainer Arlt)

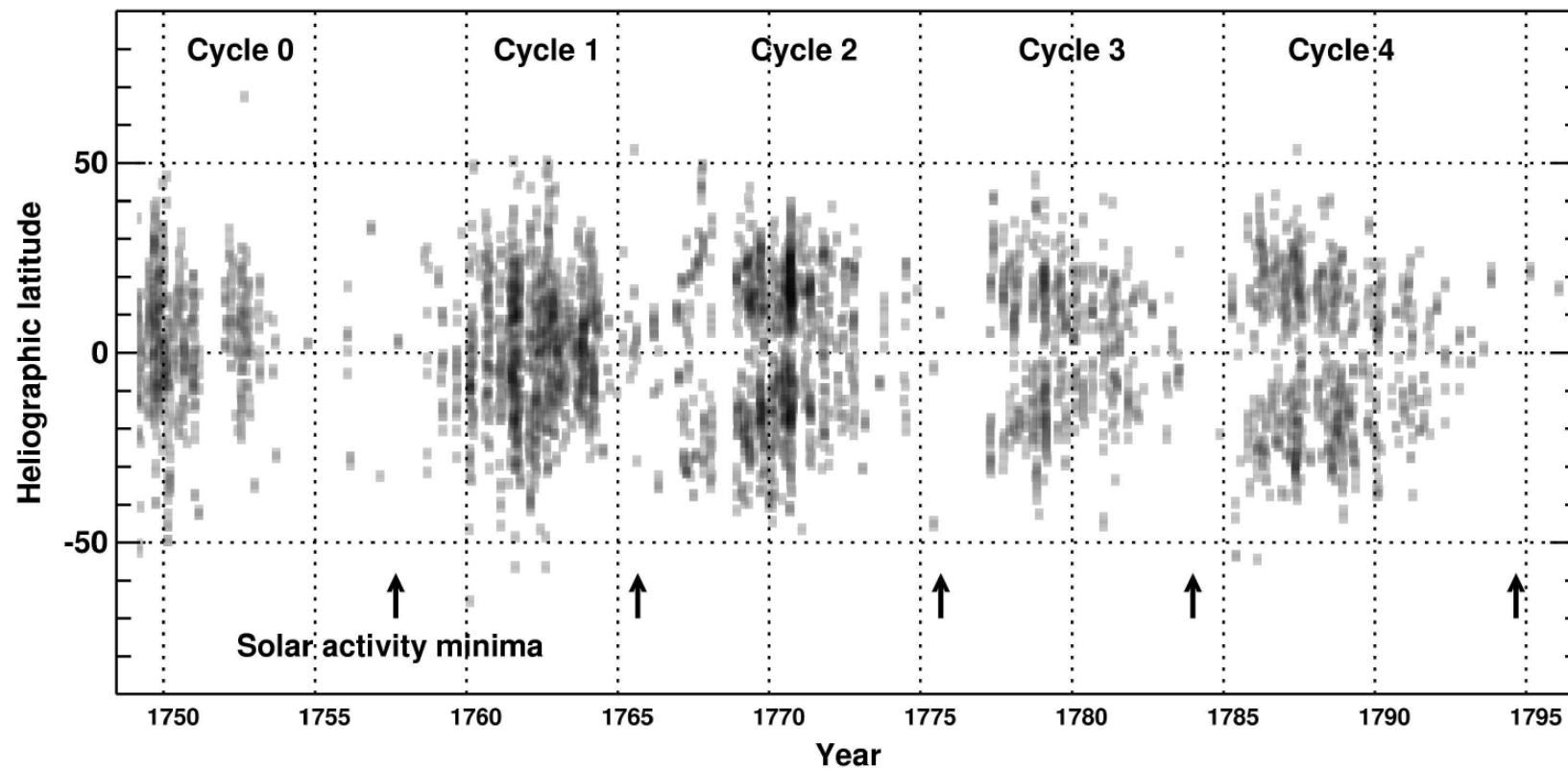
R. Arlt (2008) “Digitization of Sunspot Drawings by Staudacher in 1749 – 1796” *Solar Physics* 247, 399-410.



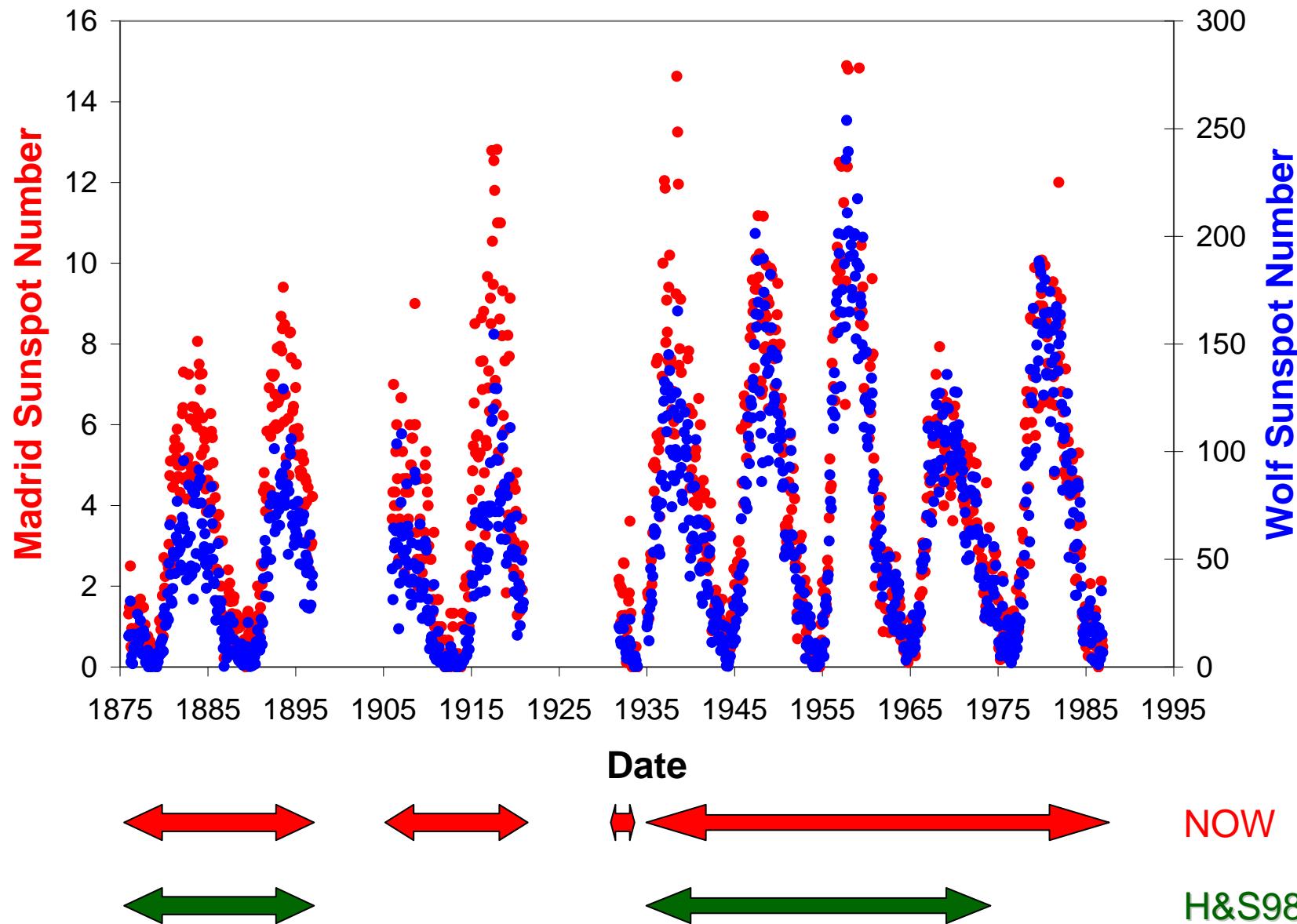
R. Arlt (2008) digitized original drawings by J.C. Staudacher made in the period of 1749 – 1796. Arlt also evaluated the usefulness of the drawings for the determination of sunspot positions for future studies.

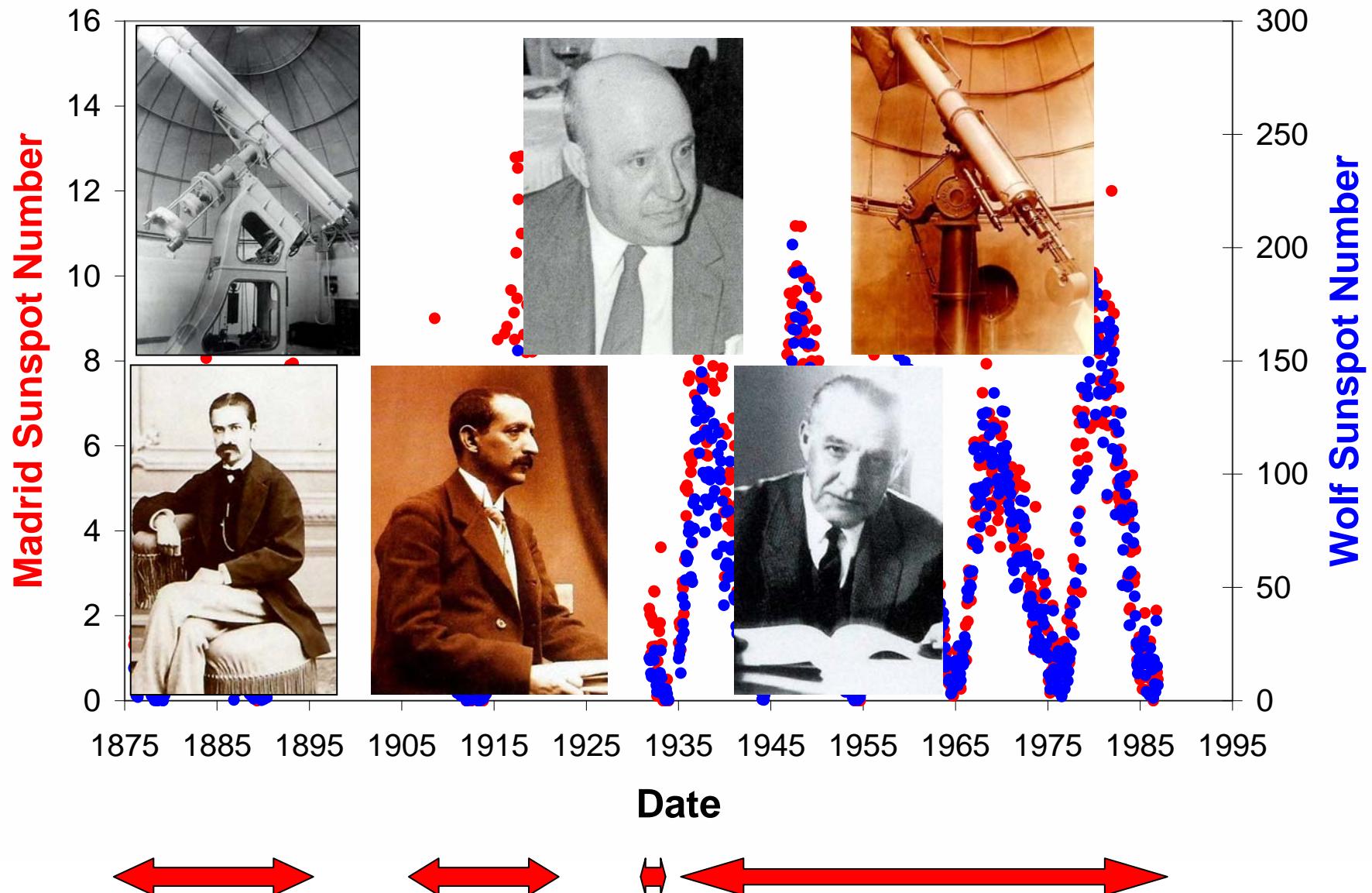


R. Arlt (2009) “The Butterfly Diagram in the Eighteenth Century” *Solar Physics* 255(1), 143-153, DOI: 10.1007/s11207-008-9306-5

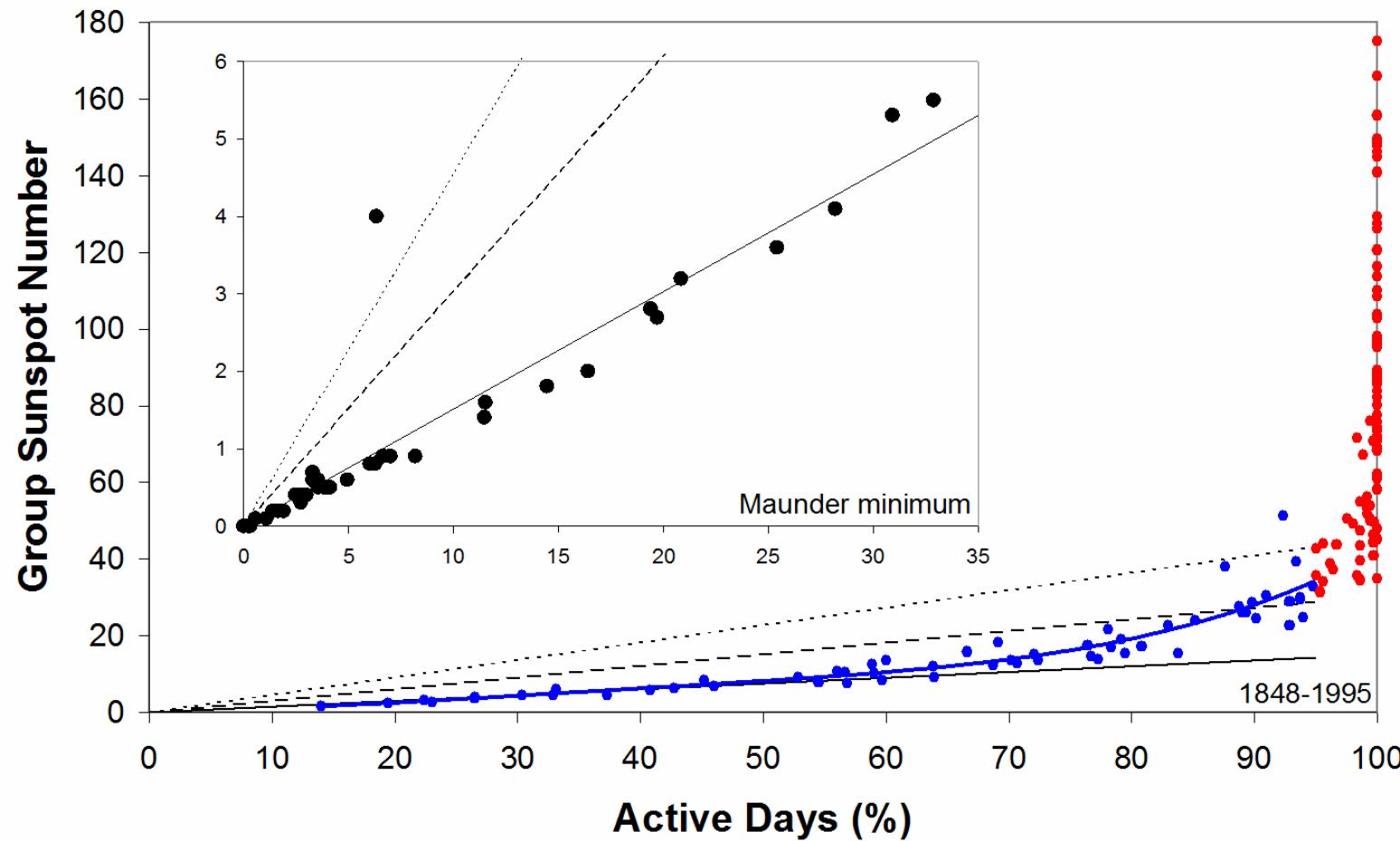


### 3.3. Madrid SSN: 110 years of solar observations

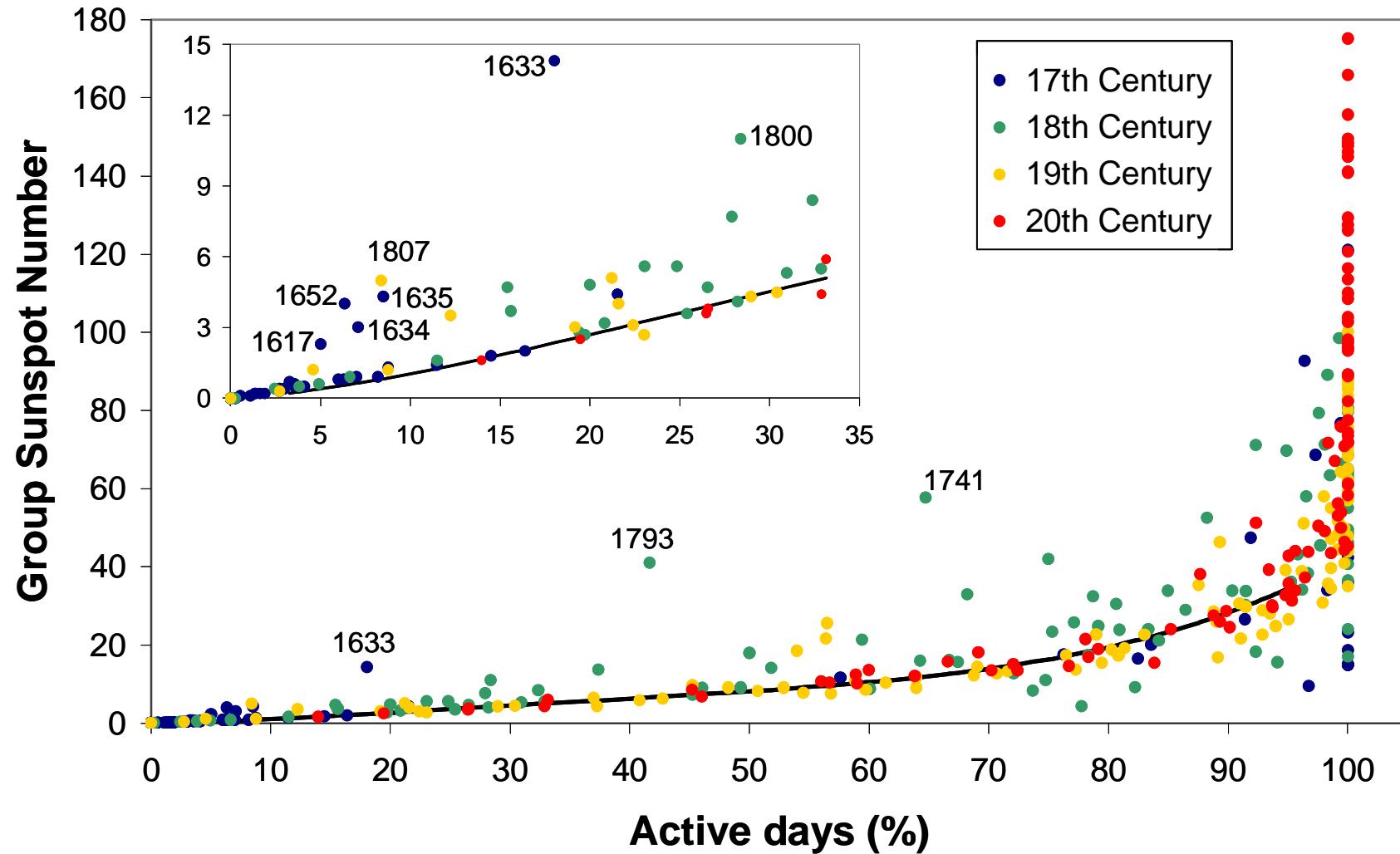




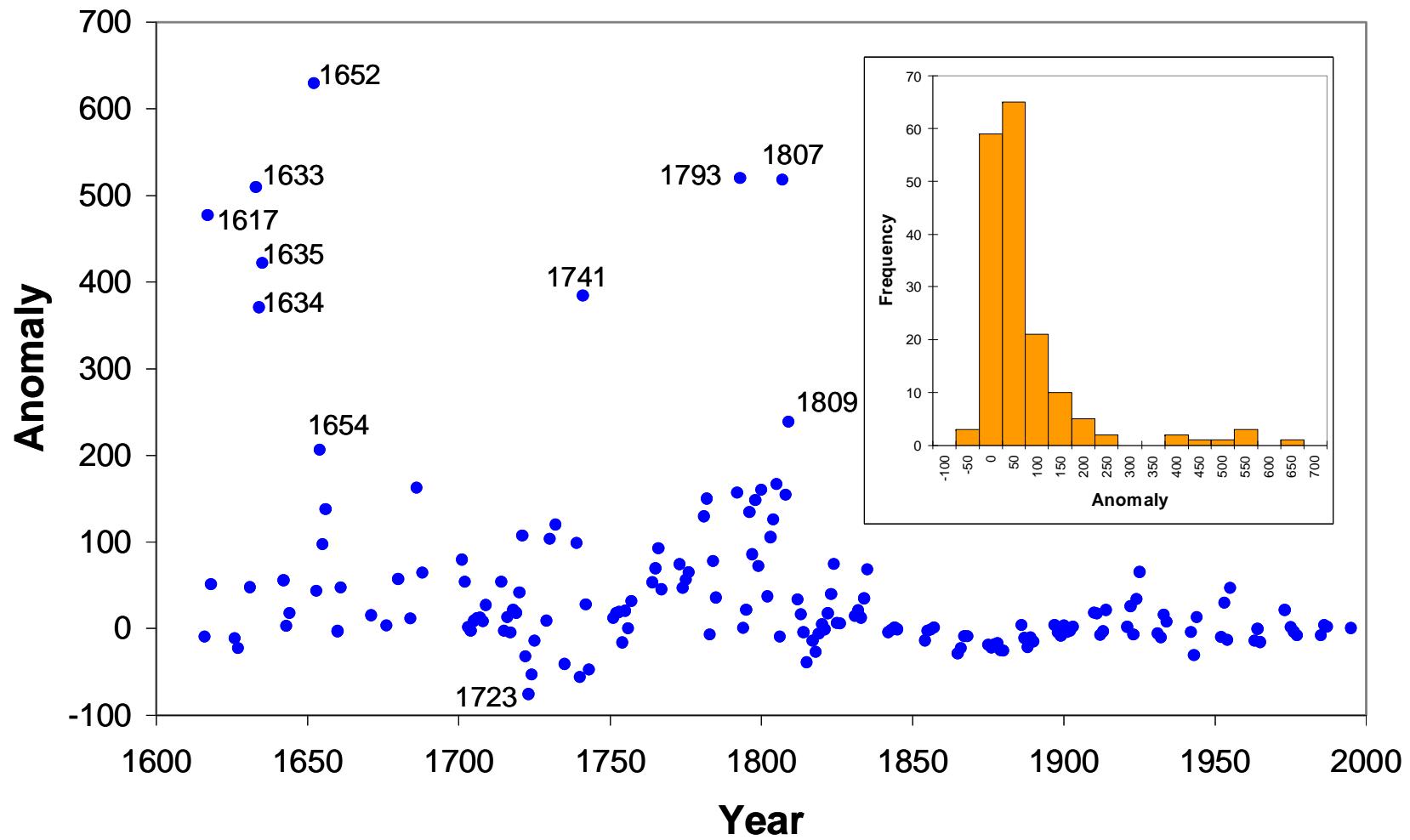
## 4. Quality Control: an example



Relationship between GSN and AD for 1848-1995 from Hoyt and Schatten (1998). Polynomial fit (order 4) is showed for AD<95% (blue line and points). Graphic inserted shows the same relationship during the Maunder minimum. Black lines represent the theoretical values for an average observer with 1 (continuous), 2 (dashed), and 3 (dotted) groups for each active day.



Relationship between GSN and AD for all available data from HS98. Black line is the polynomial fit of Figure 1. The inset presents an enlarged version but restricted to values AD<35%

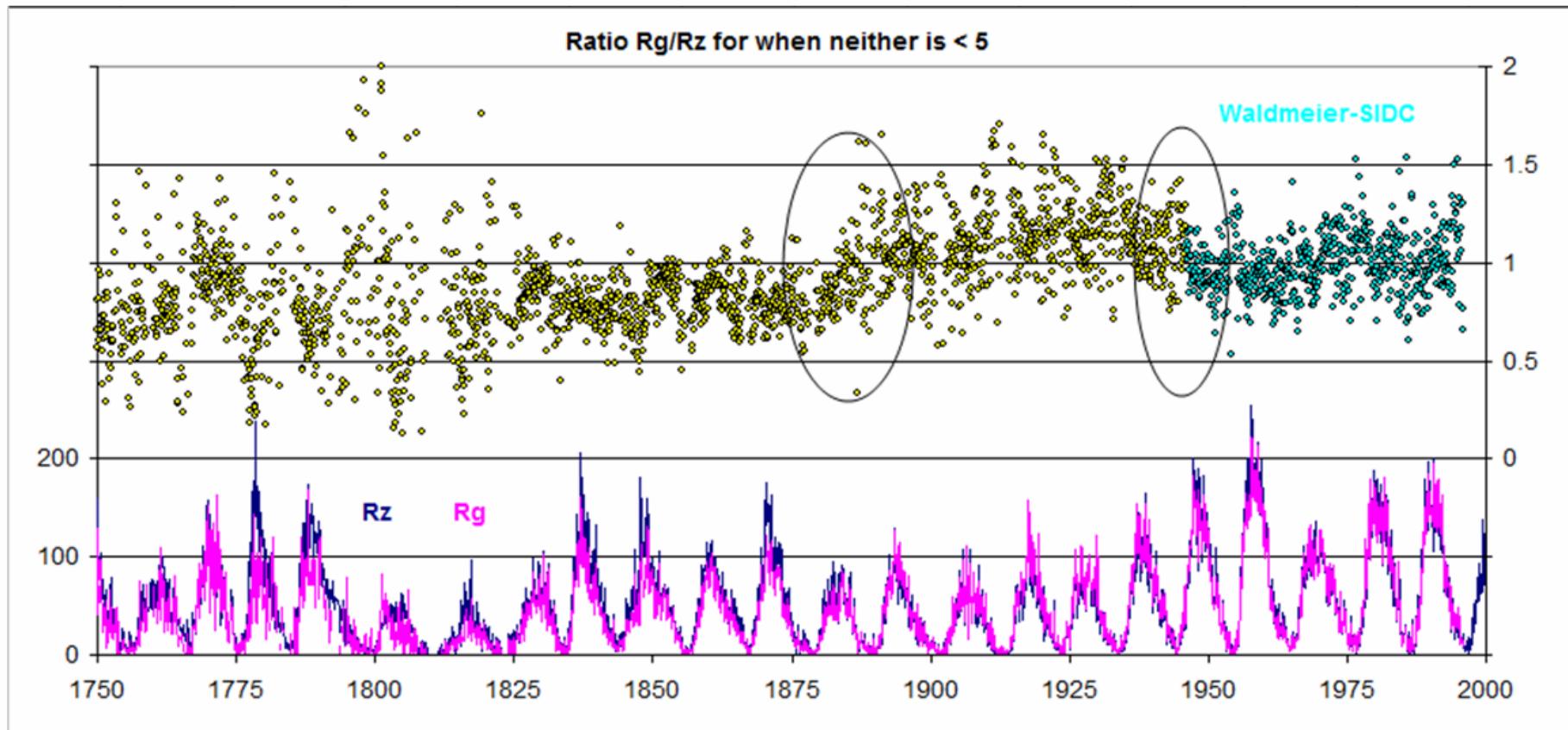


Anomaly values for GSN (blue dots). Graphic inserted shows the distribution of these values.

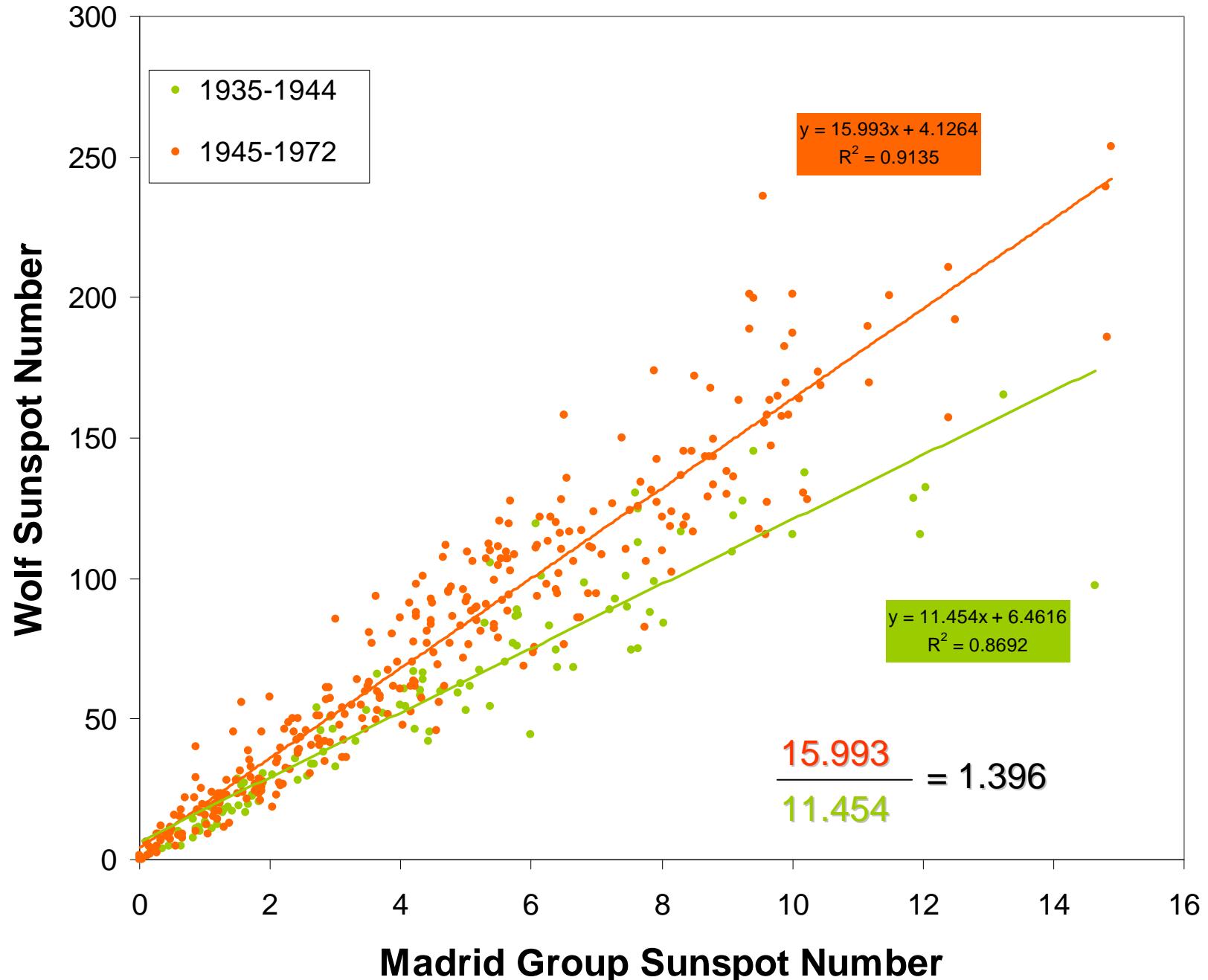
## 5. Leif's breakpoints

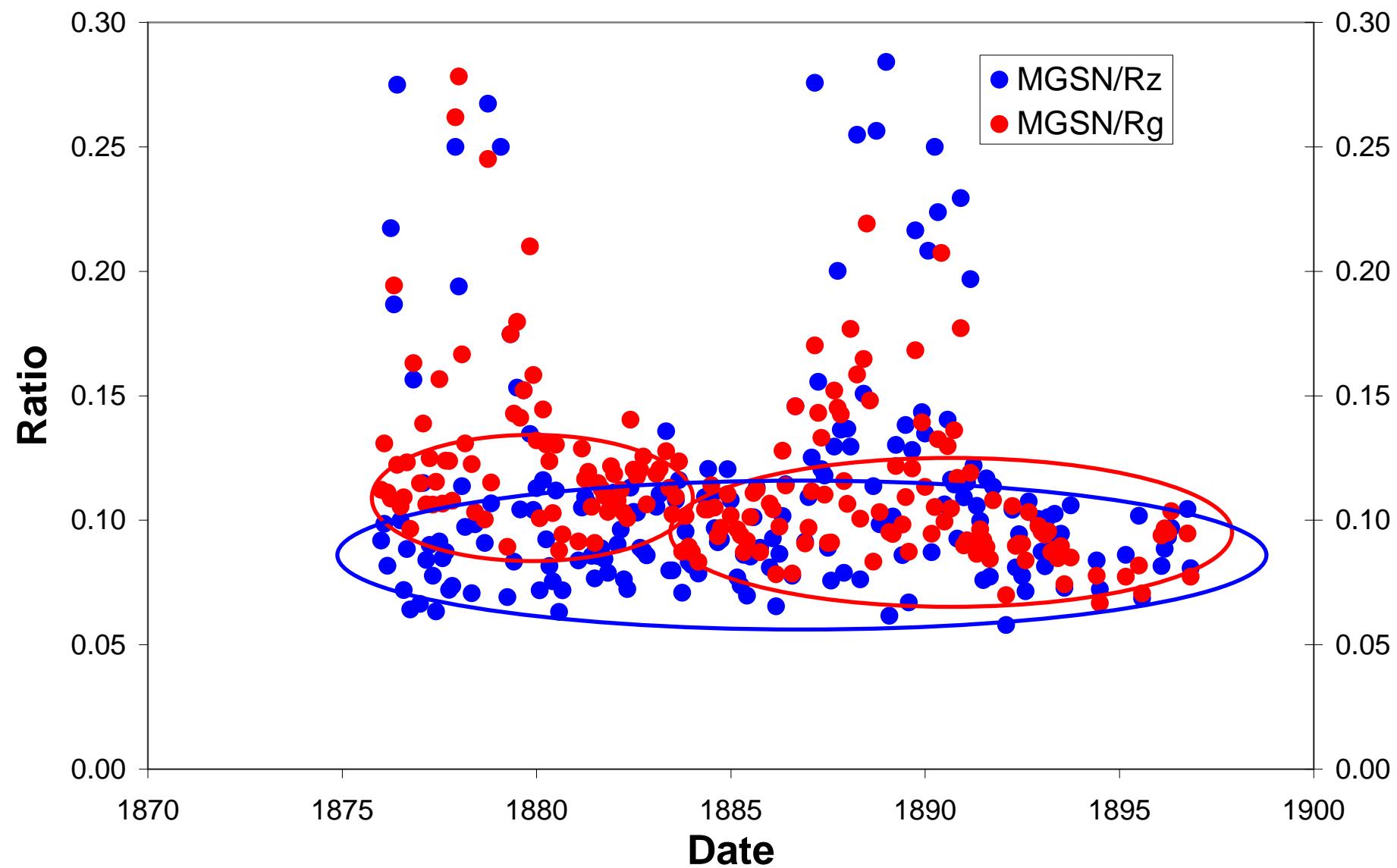
Cliver, Svalgaard & Schatten (2011)  
IUGG XXV General Assembly, IAGA  
Symposium A12.2

# The Ratio Group/Zurich SSN has Two Significant Discontinuities



At ~1946 (After Max Waldmeier took over) and at ~1885





## 6. Some additional proposals

- ✓ We need a archive or library which preserve all the historical material related with sunspots records (originals [old books, journals or manuscripts], paper copies, photocopies, digital copies, etc.).
- ✓ I think that it is possible to construct a hemispheric sunspot number using old sunspot drawings prior to RGO epoch (Staudacher, Schwabe, Peters, de la Rue, Carrington, etc..).

