



# From Galileo through the Maunder Minimum

**J.M. Vaquero & E.W. Cliver**

**4th SSN Workshop**

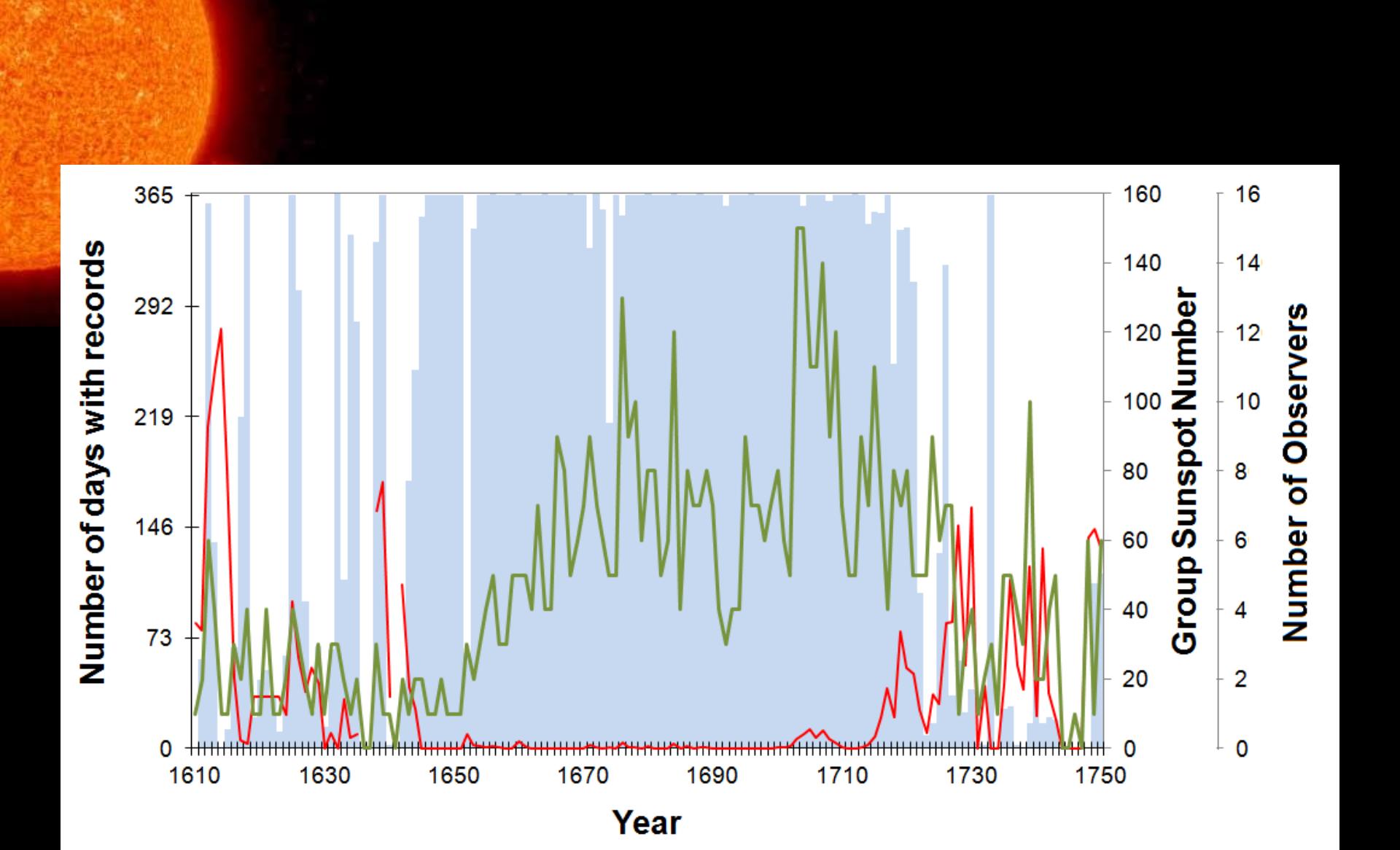
**Locarno 2014**

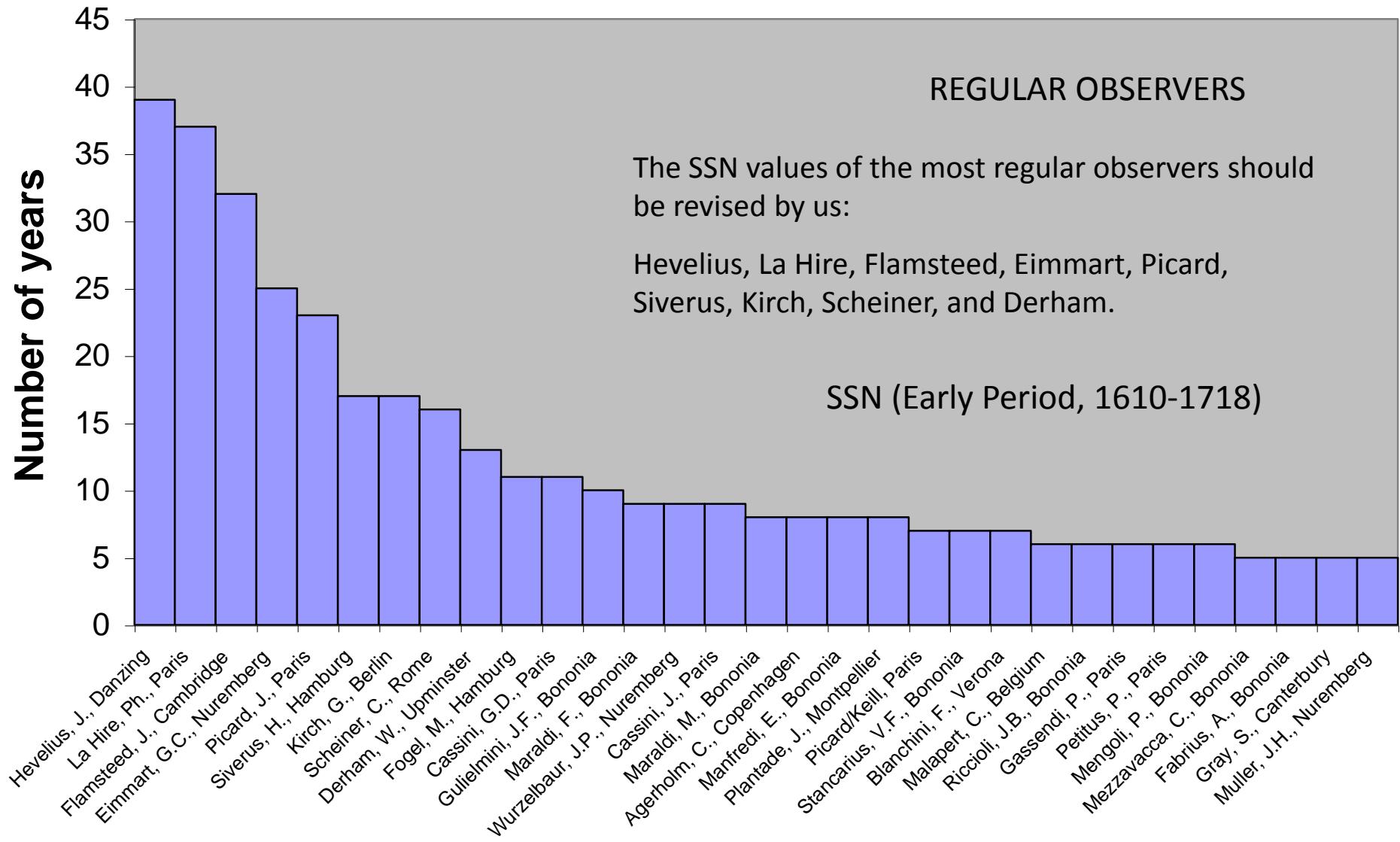


# Index

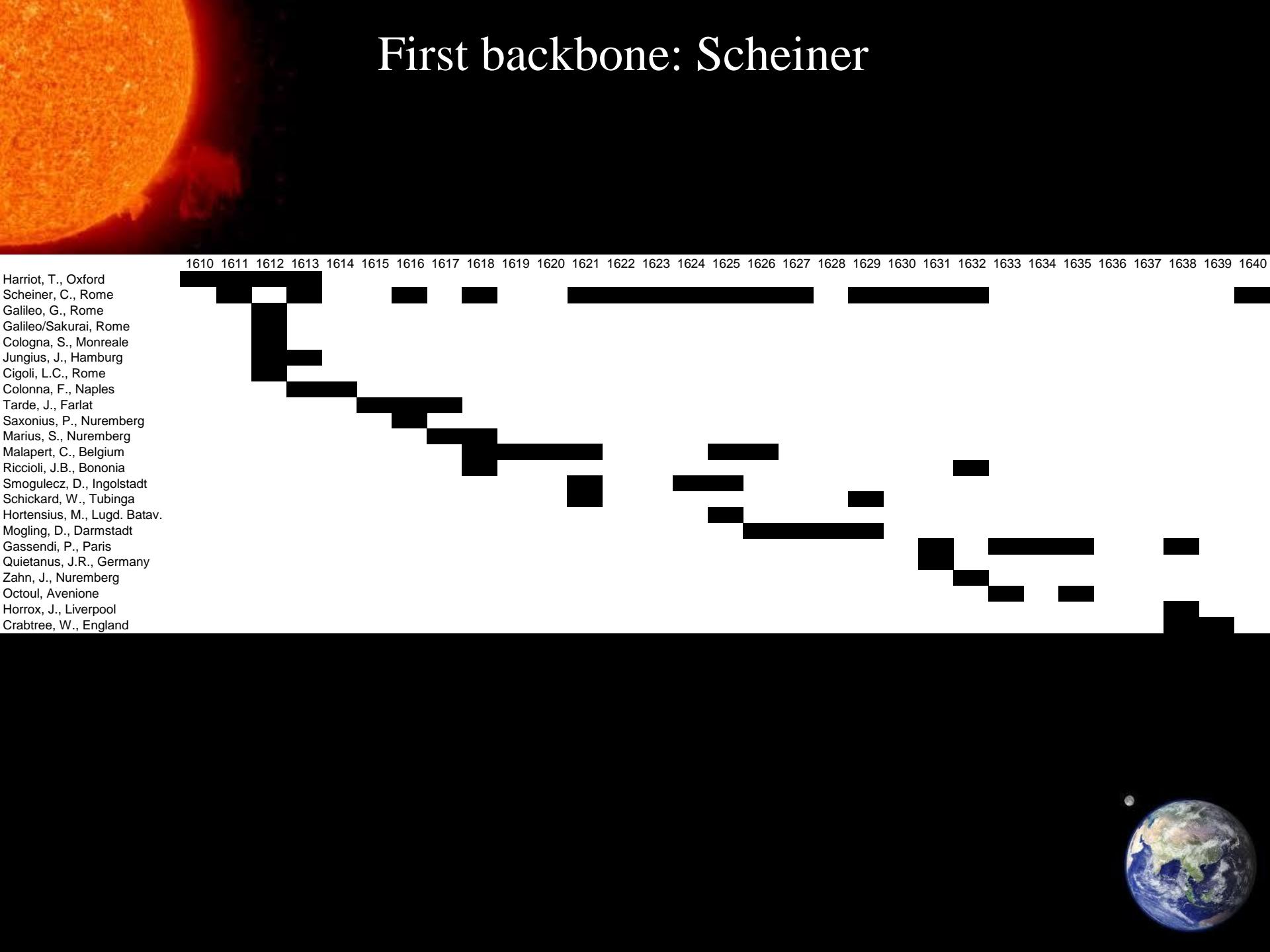
- 1) Is it possible to use the methodology of backbones in the early part of the SSN series (1610-1800 aprox.)?
- 2) The case of the Bion's sunspot (1672).
- 3) A “modern” example of solar meridian altitude measurements.
- 4) Modifying the HS98 data base....
- 5) Conclusions.



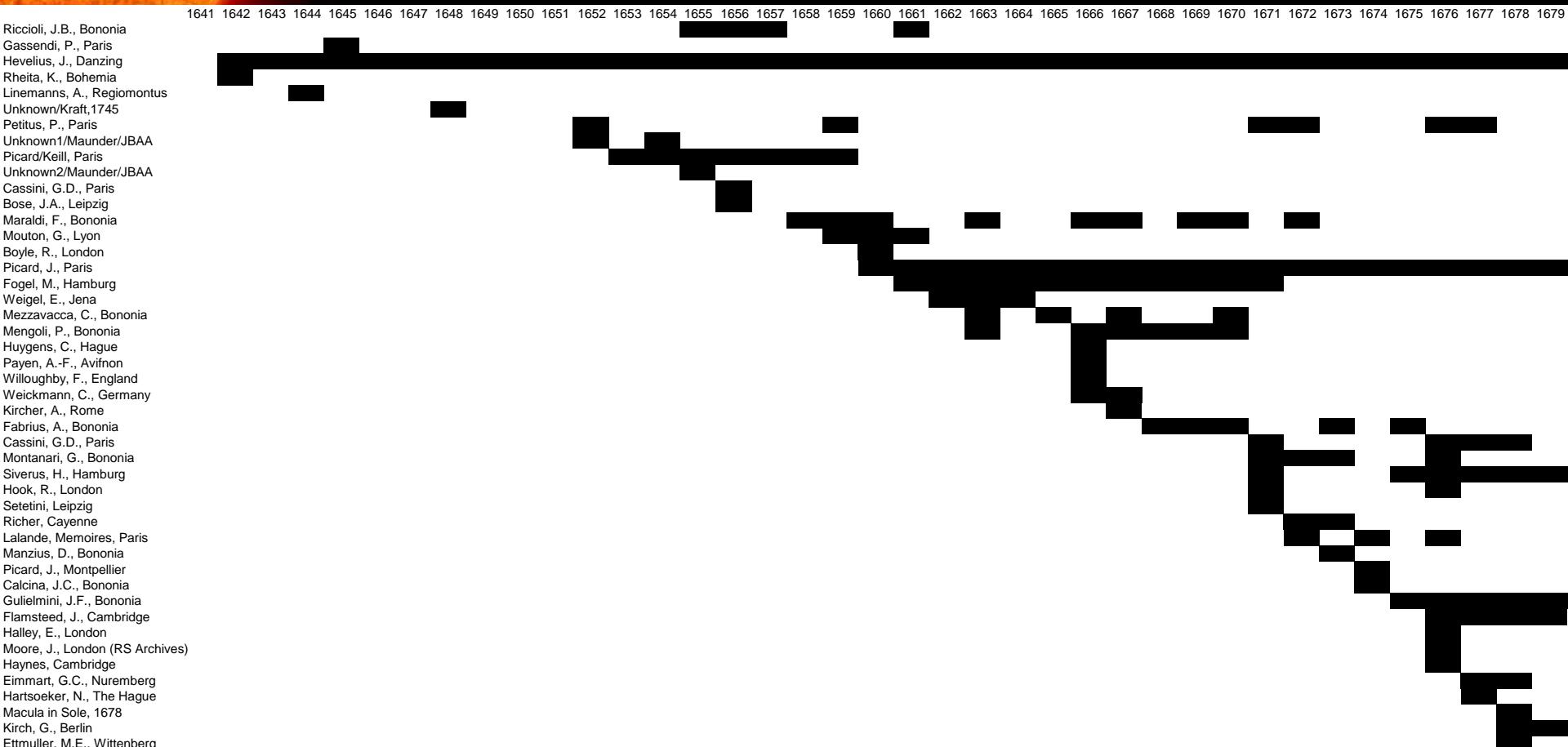




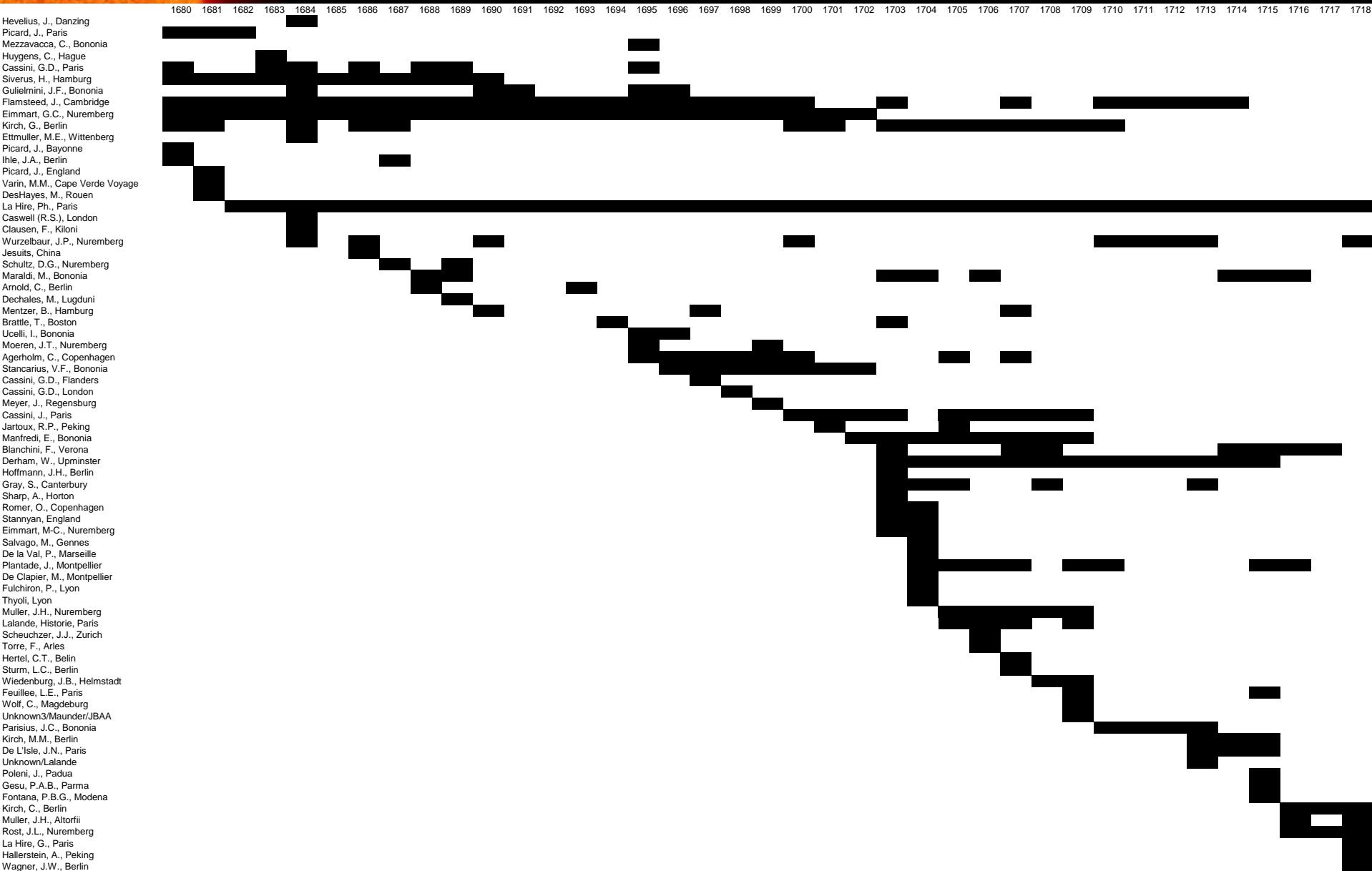
# First backbone: Scheiner



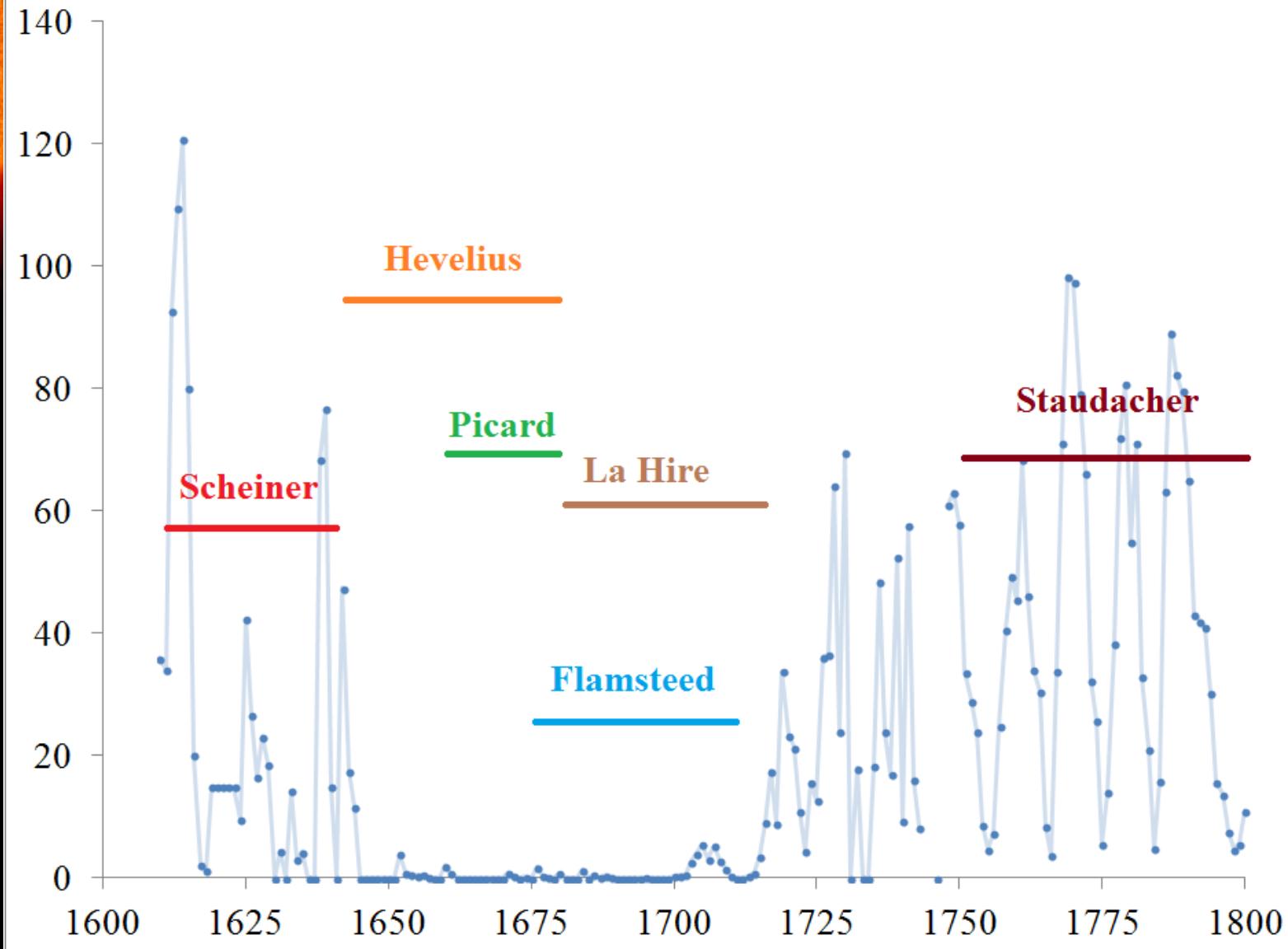
# Second backbone: Hevelius



# Third backbone: La Hire

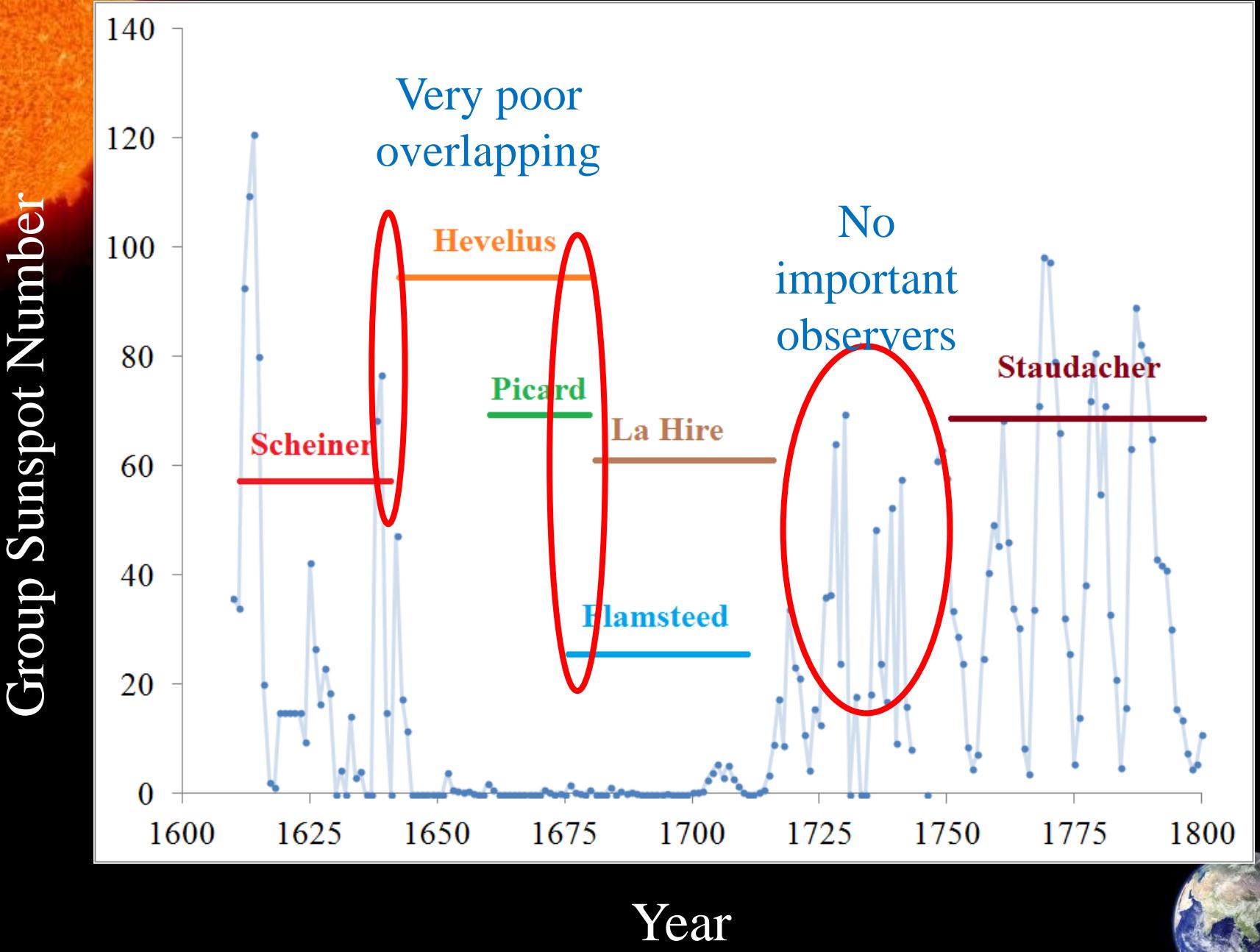


Group Sunspot Number

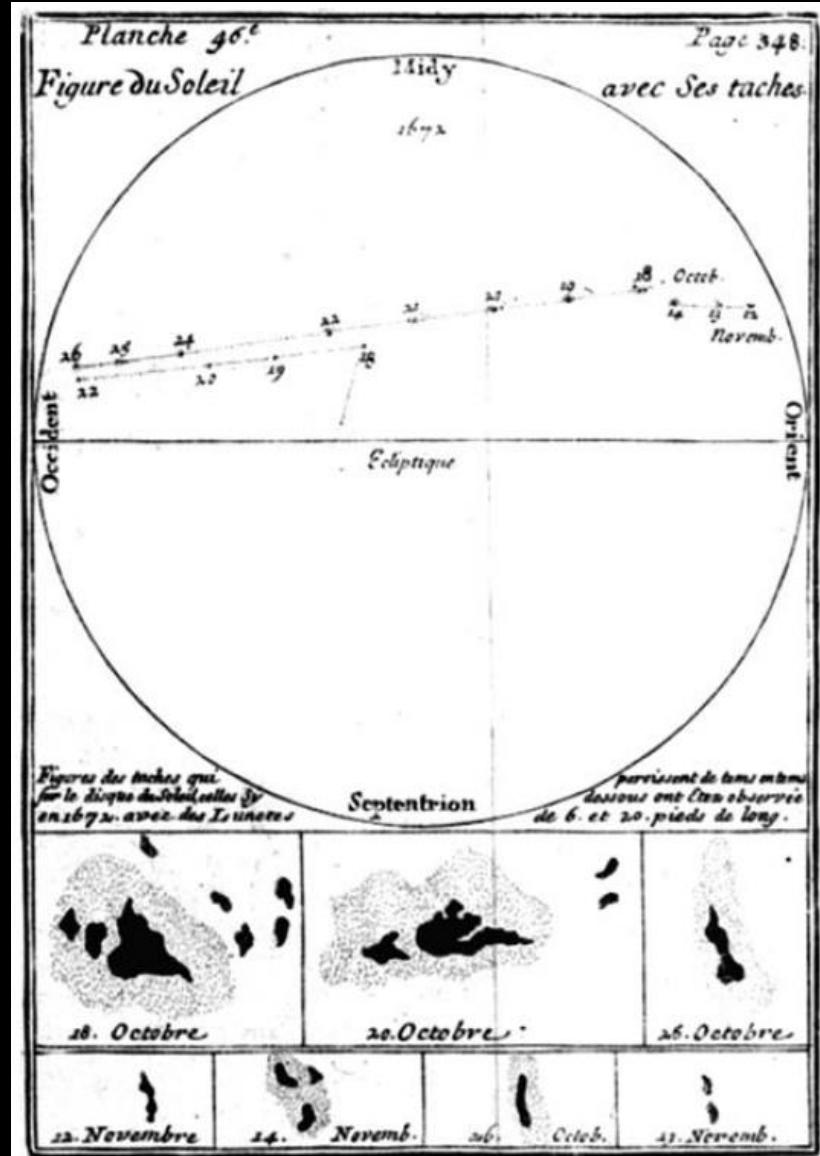


Year



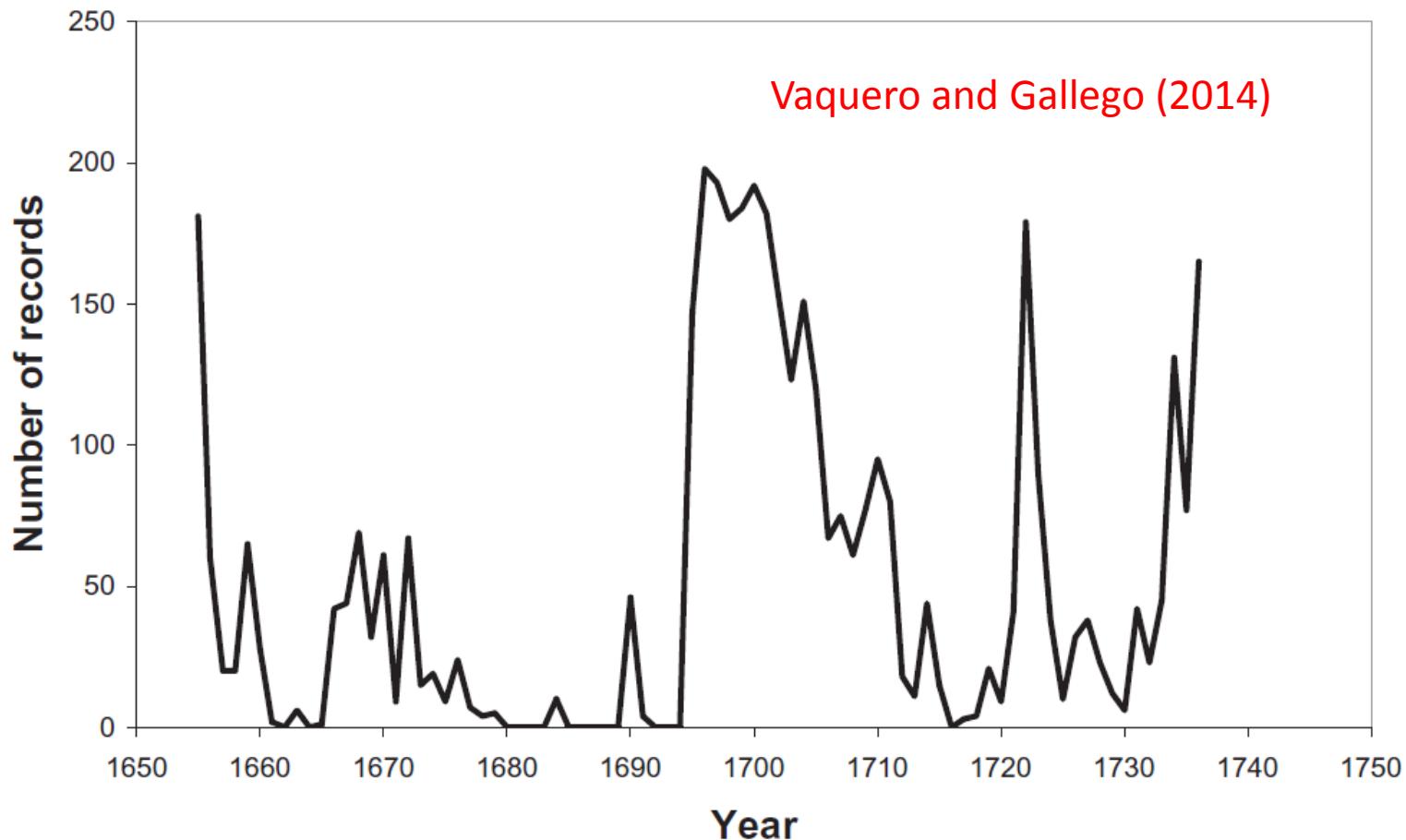


Nicholas Bion (1652-1733)  
French instrument maker



Vaquero (2007)





Number of annual solar meridian observations made in the Basilica of San Petronio recorded in [Manfredi \(1736\)](#).



# In Progress...

MACHINÆ COELESTIS LIB. III.						9
A N N O M. D C. L X.						
Mens. Dies st. n.	Altitudines Solis Meridianæ. Grad. Min. Sec.	Quo Instrumento	Quā Tempestate.	Quā Diligentia.	NOTANDA.	
Febr. 23 ♂	25 48 50	Quad. Az.	Cælo perquam sereno diligentis.		Bina macula lulum decreverant	
24 ♂			Hor. 2 p. m.	Macula parv.	Macula maj. minor erat altera vero evanuerat, nisi facula dilutissima & Umbra conspecta.	Sol in Piscibus,
26 ♀				Hor. 12.15 m.		
29 ☽					In Sole non	
Martii 1 ♂		Quad. Az.			Sol omnino purus	
	29 35 35		Cælo admodum sereno diligentis.		Nihil pariter	
					Nil macularum	
	32 20 50		Cælo perquam sereno exactissime		Nulla macula	
	32 21 0					
12 ♀	32 44 40	Quad. Az.	Cælo subnubilo	circiter	Sol purus apparet.	
Martii 14 ☽	33 31 40		Cælo subnubilo			
16 ♂	34 18 0		Cælo sudo	accuratisimè	Macula cum 2 minoribus circa Horiz. Ortiv. conspeta, quas die 14 vel 13 Sole intrasse puto.	
Martii 17 ☽						

Sunspot record with measurement of solar meridian altitude

Sunspot record without measurement of solar meridian altitude

No sunspot record with measurement of solar meridian altitude

No sunspot record without measurement of solar meridian altitude

No information about sunspot but he measured solar meridian altitude

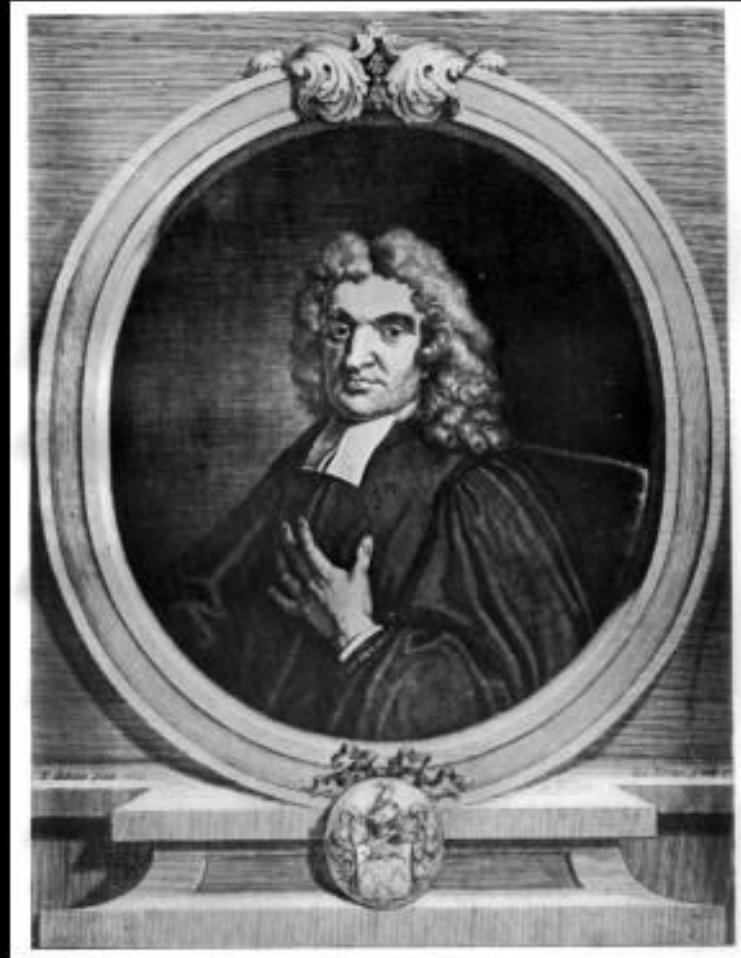
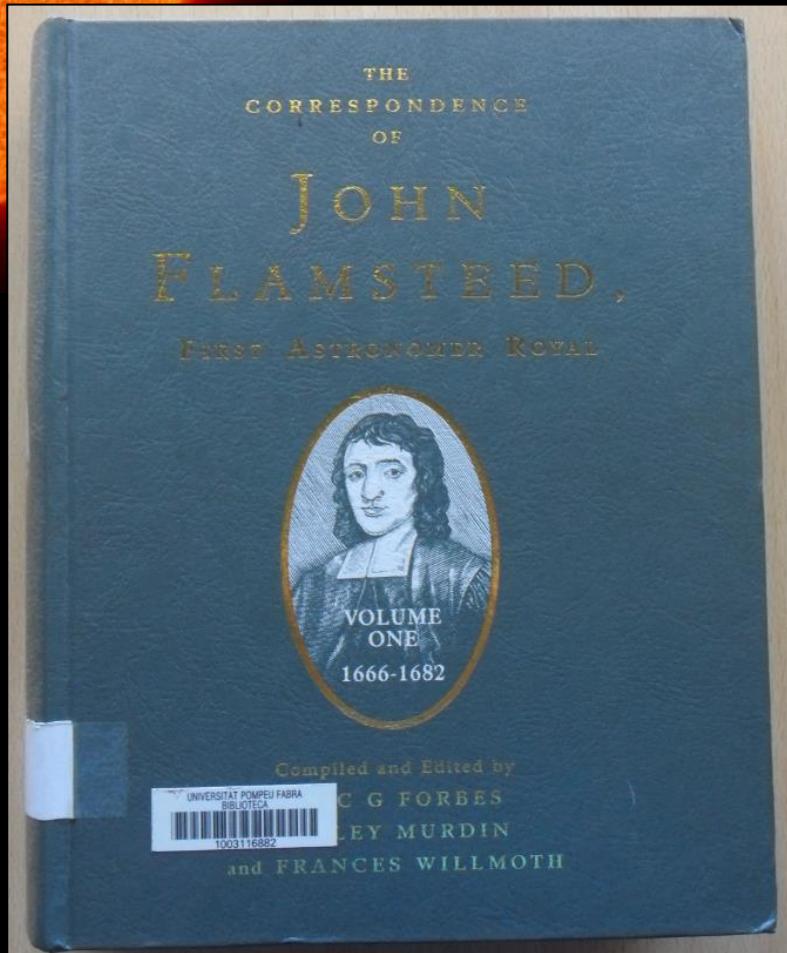


Be careful!!!

Sunspot records are not associated with measurement of solar meridian altitude!!!



# In Progress...



We are revising the sunspot record that appear in the “Correspondence” of Flamsteed. Hoyt and Schatten (1998) did not consult this historical source because was published when they were finishing their work.  
He mentions very few times the absence of spots in the Sun in exact dates!

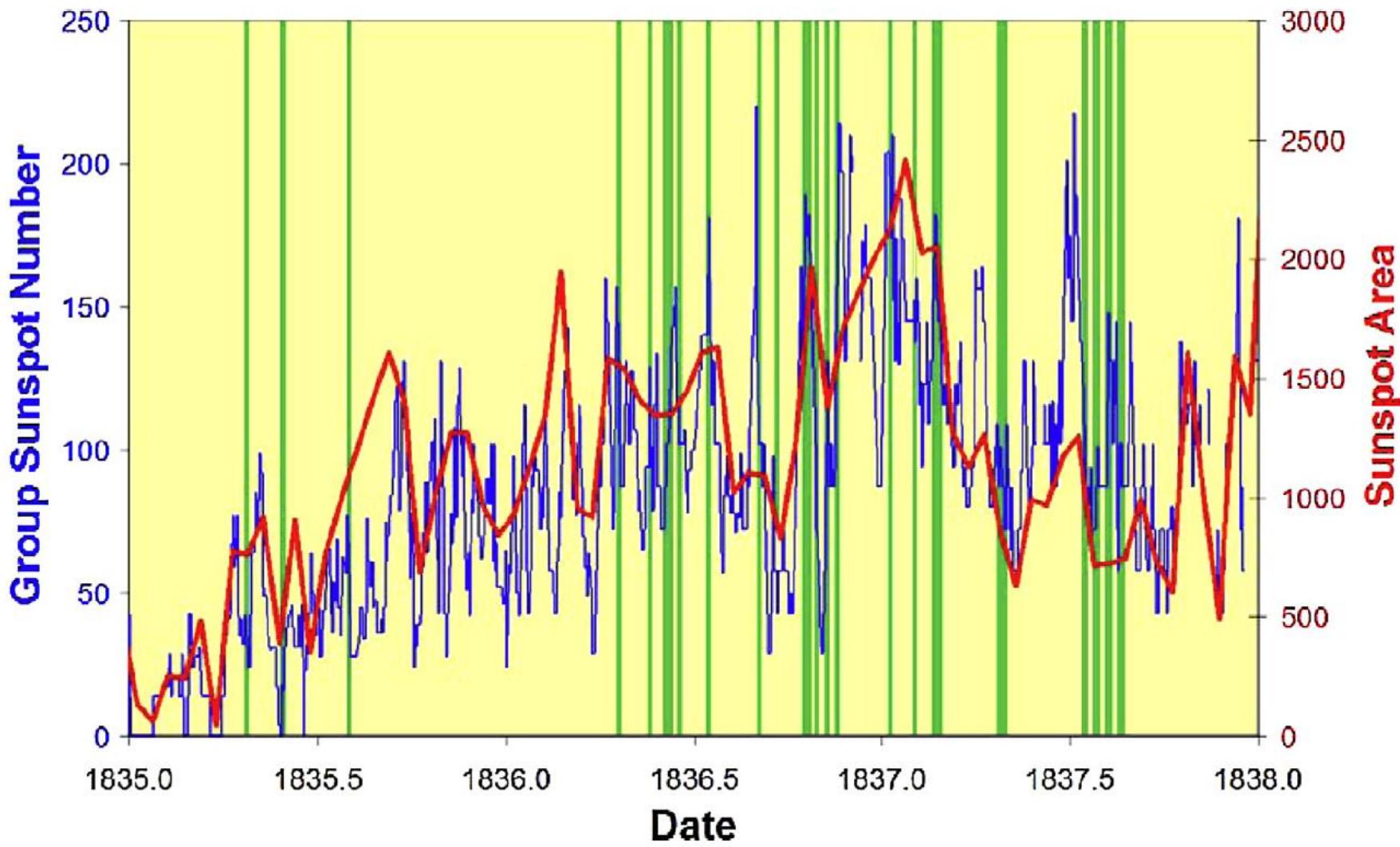


# A “modern” example...



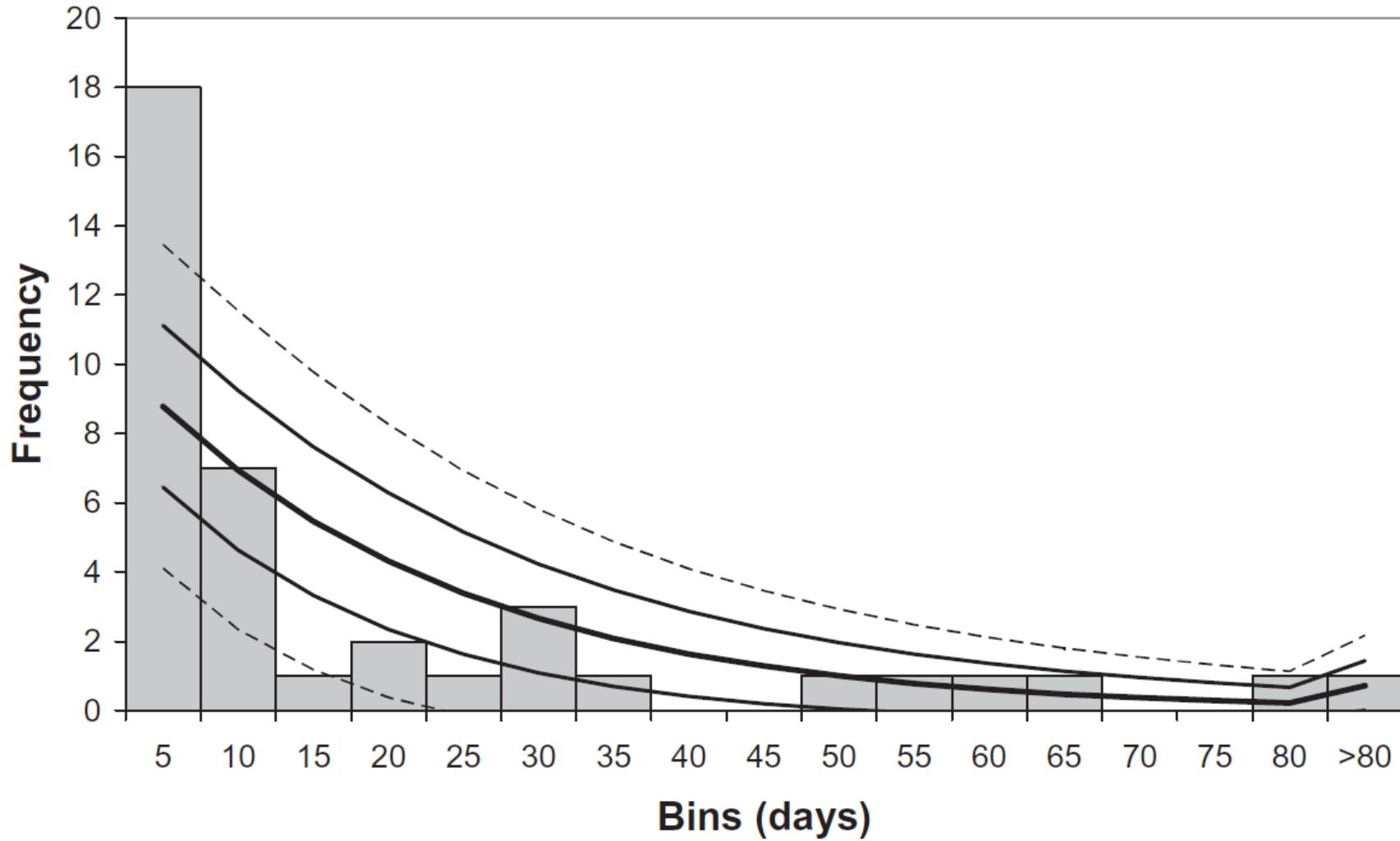
The meridian solar observations that have been consulted were made with the meridian telescope ( $D = 0.125$  m and  $f = 3.05$  m) by Thomas Jones, following the meridian telescopes installed in Greenwich and Kensington.

It was installed in ROA in 1833, being in use until 1862.



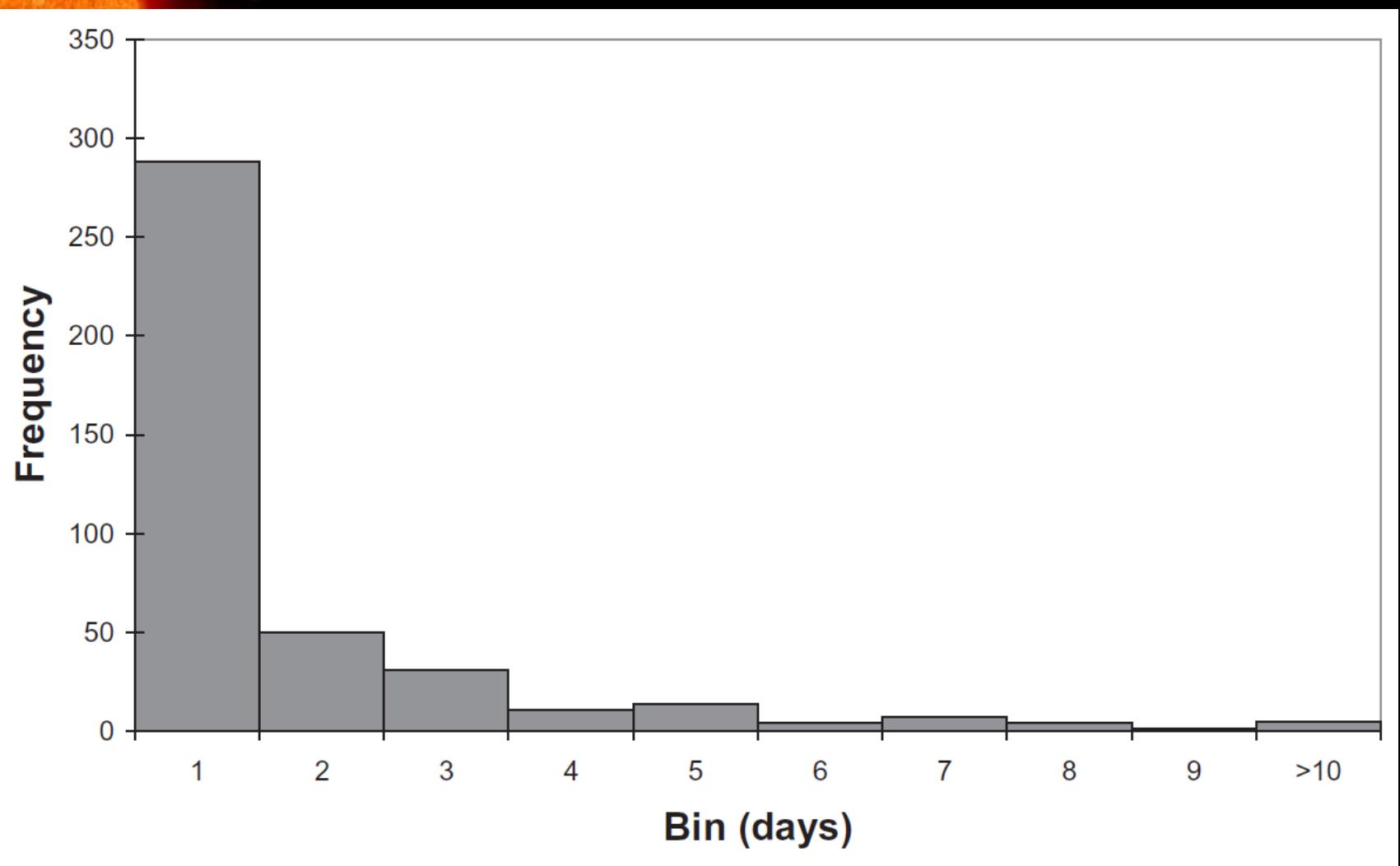
A comparison between sunspot reports in the meridian solar observations of ROA (green) and two solar indices in the same period: (red) fortnightly values of sunspot area provided by De la Rue et al. (1869) and (blue) daily values of Group Sunspot Number provided by Hoyt and Schatten (1998).





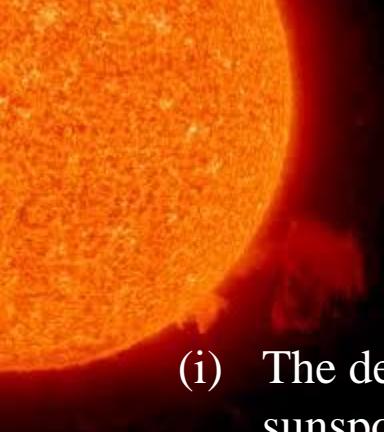
Frequency histogram on the number of cases as a function of the lapse time between sunspot reports. The mean value (thick line) are also shown with the corresponding values associated to  $\pm 1\sigma$  (thin line) and  $\pm 2\sigma$  (dashed line) intervals obtained with a Monte-Carlo simulation that have generated 200,000 random time series (constrained with a similar amount of sunspot reports).





Frequency histogram on the days between solar meridian observations made in the ROA from 25 April 1835–23 August 1837.





## In Brief...

- (i) The descriptions given (in haso.unex.es) should be considered as a list of notable sunspots.
- (ii) The sunspot records appear only for the years 1835–1837. Manuscripts of the observations of 1833–1834 and 1838–1840 contain no observations of sunspots.
- (iii) Records of sunspots only appear in the manuscripts of the observations. These records do not appear in the printed version of the meridian observations.
- (iv) There is even an incorrect record of no sunspot (although we do not know the origin of this error).
- (v) There is not a clear relationship between the timing of recorded spots and usual solar activity indices (sunspot area and number).
- (vi) There are no notable periodicities in these observations.

In general, solar meridian observations should be used with extreme caution to evaluate past solar activity.





# Modifications of the HS98 data base

- (i) We have eliminated all the observations made with the camera obscura of San Petronio Basilica. (In total, 4204 “zero” values removed)
- (ii) We have eliminated all the record of astronomer that observed during 365 days per year (or almost). (In total 75·365 “zero” values removed, aprox. 27000 values)
- (iii) We have included some “new” observers.
- (iv) We have included all the (few) corrections to the HS98 data base published since 1998.

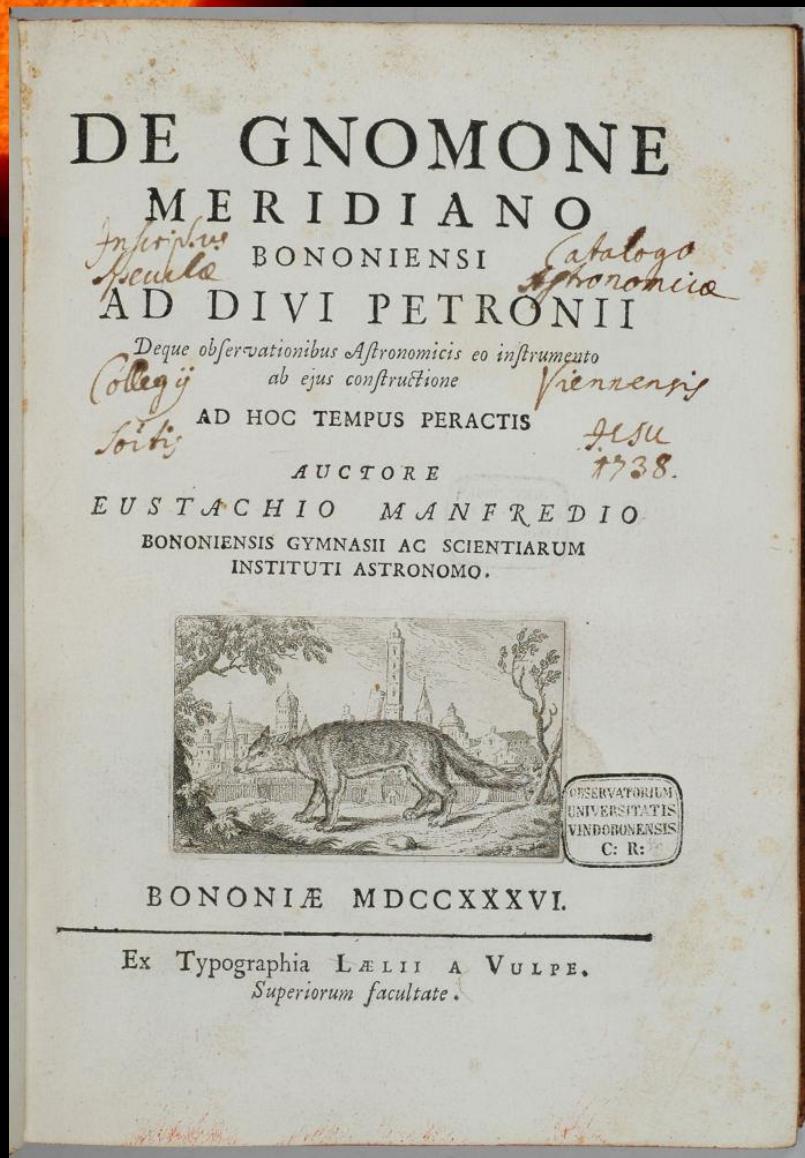
Original HS98 database  
58 788 values



Modified HS98 database  
32 339 values



(i) We have eliminated all the observations made with the camera obscura of San Petronio Basilica.



(ii) We have eliminated all the record of astronomer that observed during 365 days per year (or almost).

NUMBER OF SUNSPOT GROUPS FOR THE YEAR: 1675 AS OBSERVED BY: SIVERUS, H., HAMBURG													
Day	Jan	Feb	Mar	Apr	May	Jun	Jul	Aug	Sep	Oct	Nov	Dec	
1	0	0	0	0	0	0	0	0	0	0	0	0	0
2	0	0	0	0	0	0	0	0	0	0	0	0	0
3	0	0	0	0	0	0	0	0	0	0	0	0	0
4	0	0	0	0	0	0	0	0	0	0	0	0	0
5	0	0	0	0	0	0	0	0	0	0	0	0	0
6	0	0	0	0	0	0	0	0	0	0	0	0	0
7	0	0	0	0	0	0	0	0	0	0	0	0	0
8	0	0	0	0	0	0	0	0	0	0	0	0	0
9	0	0	0	0	0	0	0	0	0	0	0	0	0
10	0	0	0	0	0	0	0	0	0	0	0	0	0
11	0	0	0	0	0	0	0	0	0	0	0	0	0
12	0	0	0	0	0	0	0	0	0	0	0	0	0
13	0	0	0	0	0	0	0	0	0	0	0	0	0
14	0	0	0	0	0	0	0	0	0	0	0	0	0
15	0	0	0	0	0	0	0	0	0	0	0	0	0
16	0	0	0	0	0	0	0	0	0	0	0	0	0
17	0	0	0	0	0	0	0	0	0	0	0	0	0
18	0	0	0	0	0	0	0	0	0	0	0	0	0
19	0	0	0	0	0	0	0	0	0	0	0	0	0
20	0	0	0	0	0	0	0	0	0	0	0	0	0
21	0	0	0	0	0	0	0	0	0	0	0	0	0
22	0	0	0	0	0	0	0	0	0	0	0	0	0
23	0	0	0	0	0	0	0	0	0	0	0	0	0
24	0	0	0	0	0	0	0	0	0	0	0	0	0
25	0	0	0	0	0	0	0	0	0	0	0	0	0
26	0	0	0	0	0	0	0	0	0	0	0	0	0
27	0	0	0	0	0	0	0	0	0	0	0	0	0
28	0	0	0	0	0	0	0	0	0	0	0	0	0
29	0	-99	0	0	0	0	0	0	0	0	0	0	0
30	0	-99	0	0	0	0	0	0	0	0	0	0	0
31	0	-99	0	-99	0	-99	0	0	-99	0	-99	0	0
means: 0.0 0.0 0.0 0.0 0.0 0.0 0.0 0.0 0.0 0.0 0.0 0.0 0.0 0.0													

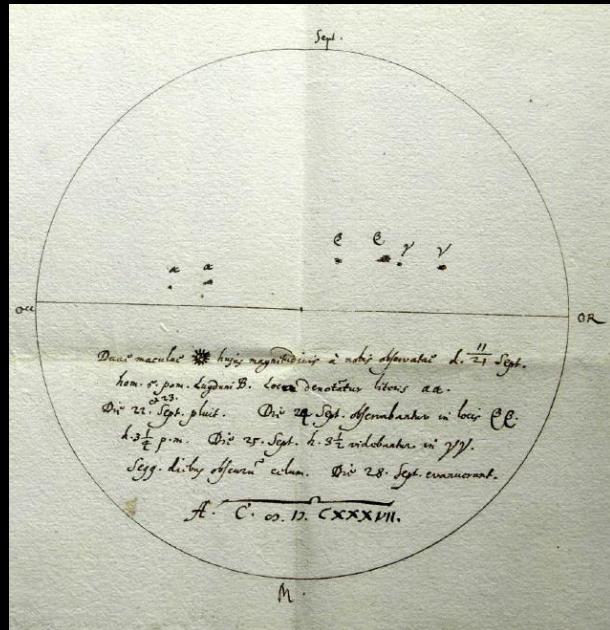
May be in Egypt  
or Cuba but ...



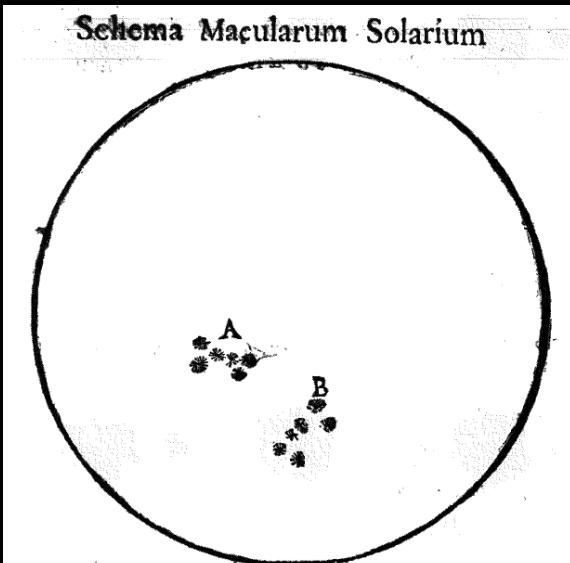
(iii) We have included some “new” observers



Andrea Argoli (Venice) 1634



Georg Marcgraf (Leiden) 1637



Conclusiones propositæ, vt evidenter suæ veritatis certitudinem sortiantur, maculas, ante aliquot dies, Anno præsenti 1640. in corpore soli obseruatas, apponere placuit: quarum altera nimis tum A. die 5. Septemb. Hora 5. post meridiem: Altera videlicet B. die 6. Septemb. Hora 11. ante meridiem sub tali formâ vt Diagramma exhibet, comparuit.

Albert Strazyc (Krakow) 1640



J. A. Alzate (Mexico) 1780s

Went to Post Office back West.

W. noon. of gulf numb. of spots in q C & stars saw. One ♀ discovered with my naked eye (but only a colored glass to save it) with the telescope appeared to be a cluster of spots exceed black & very small with a nebula; the latter yet, however, is not the part of q C, & is even a considerable arrow, with all of which was covered by 2 clouds. The noon winter has been very severe, Port Harbor quite froze up, loaded sleds driven over Charlestown ferry, &c.

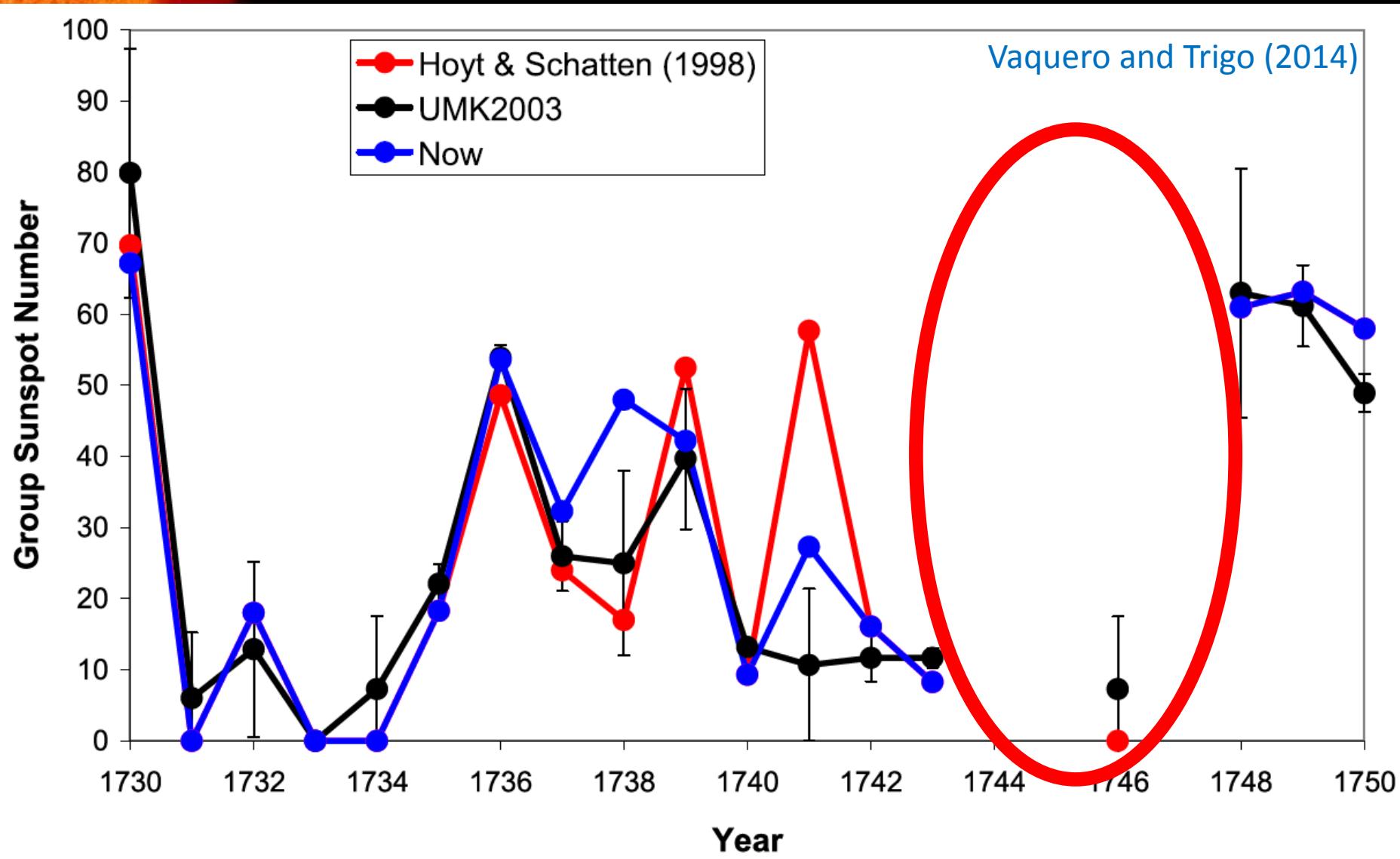
1. Show.  
2. a gift, show

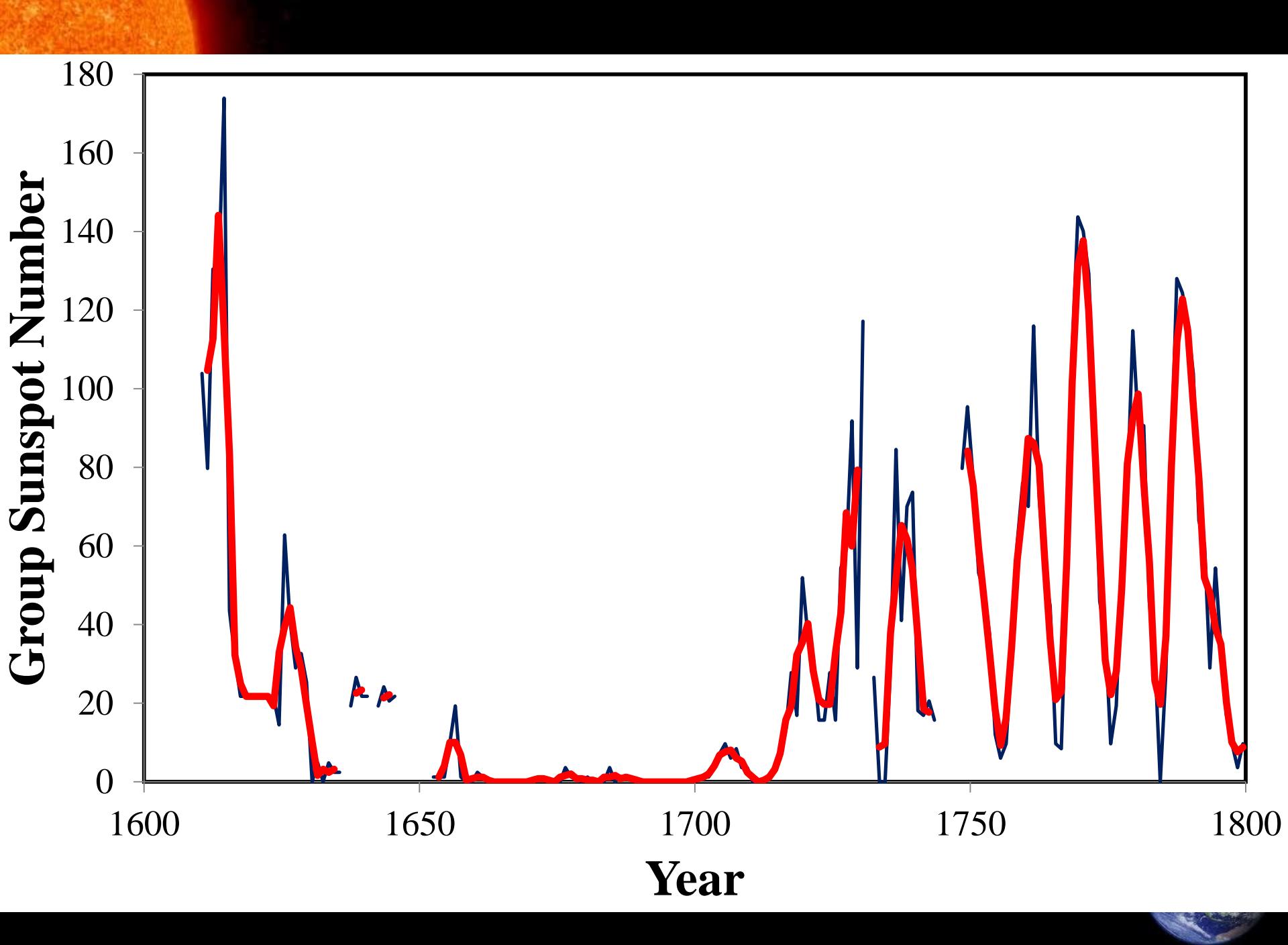
B. Glorify, warm, alt. noon had a sight of a gull spot  
" with only 4 red glints.

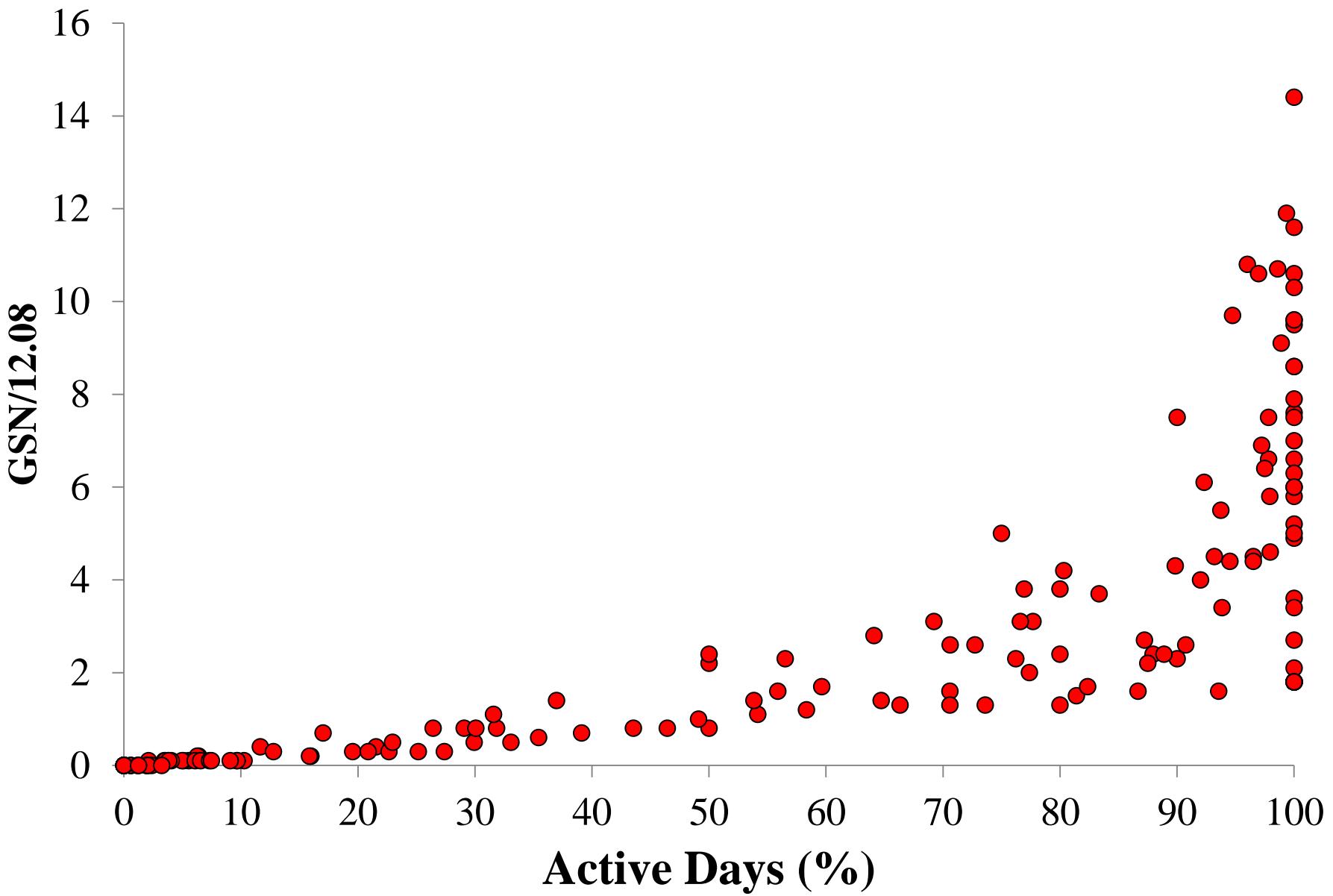
## **John Winthrop's annotated almanac (HUM 9, Box 4, Volume 2)**

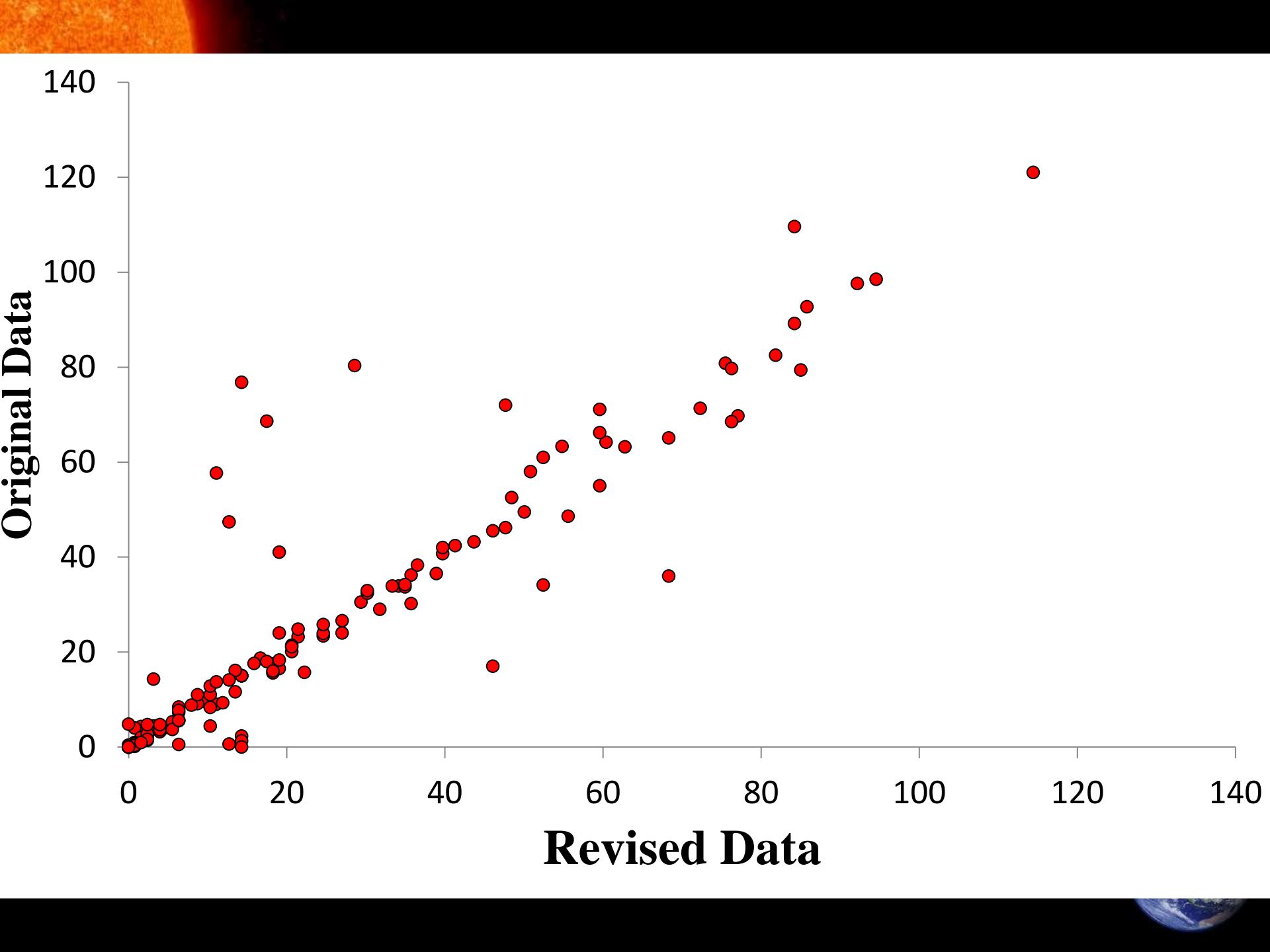


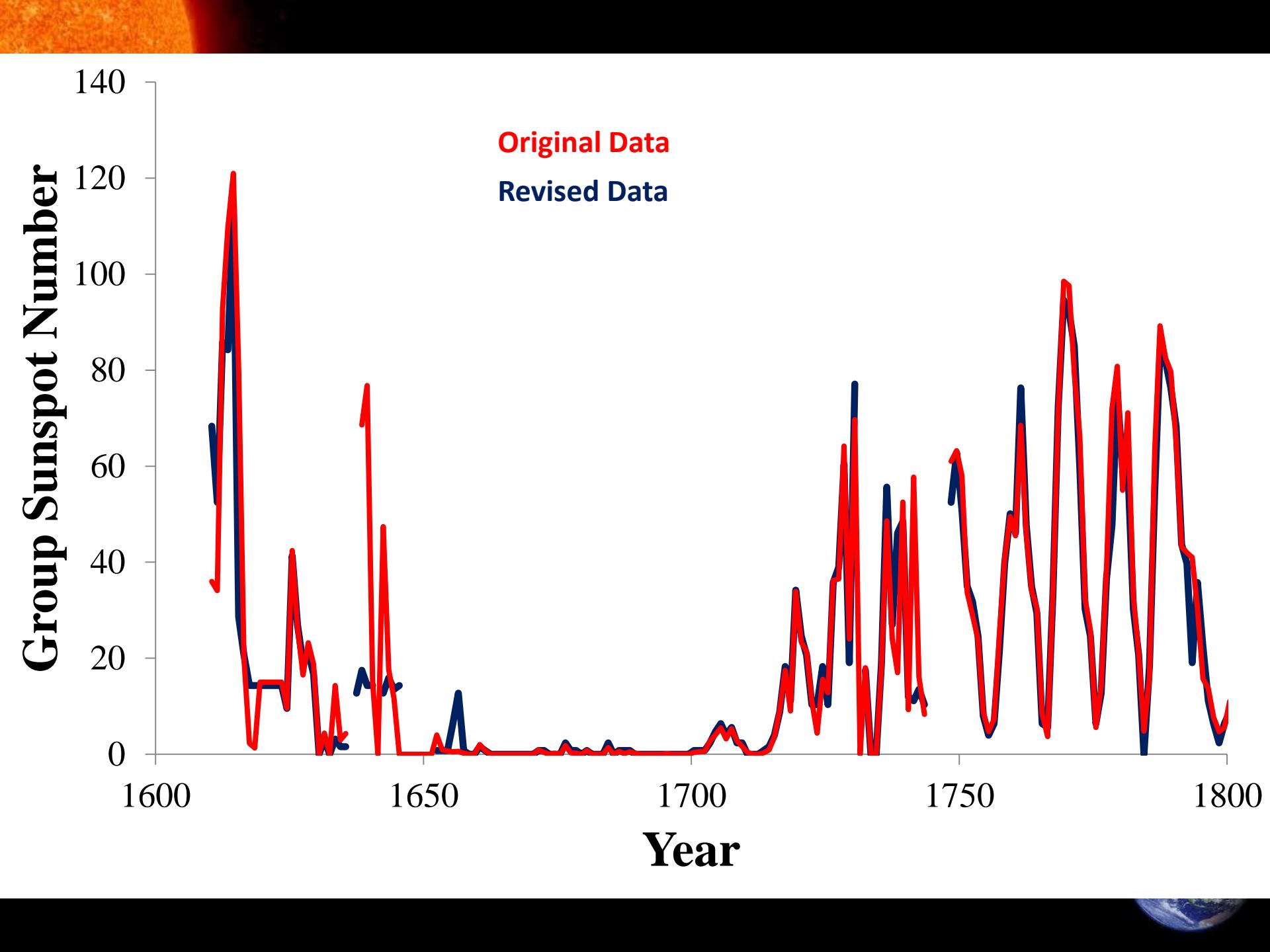
(iv) We have included all the (few) corrections to the HS98 data base published since 1998







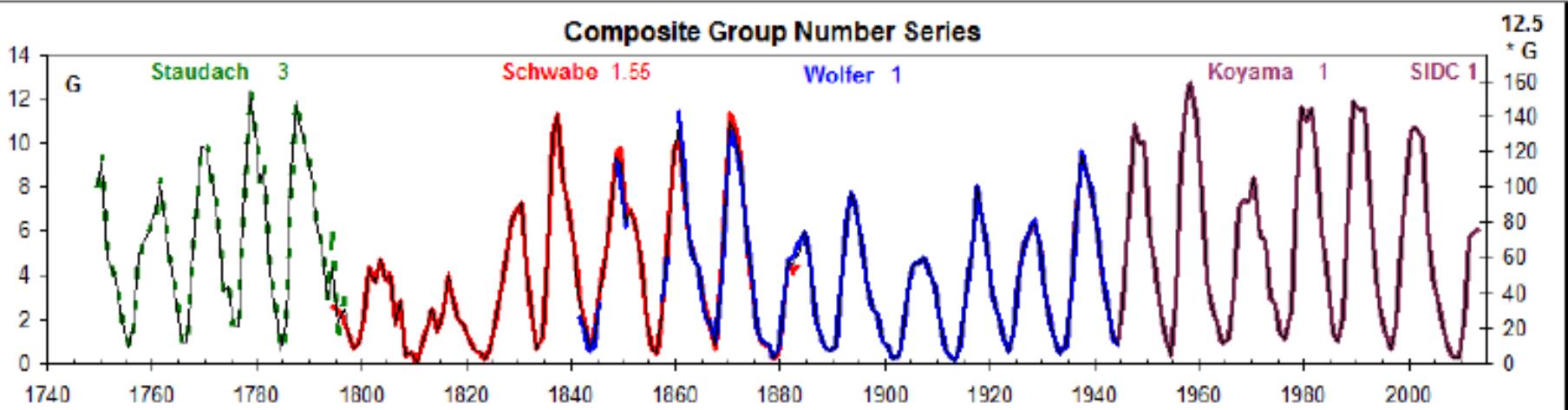




# Conclusions (first part)

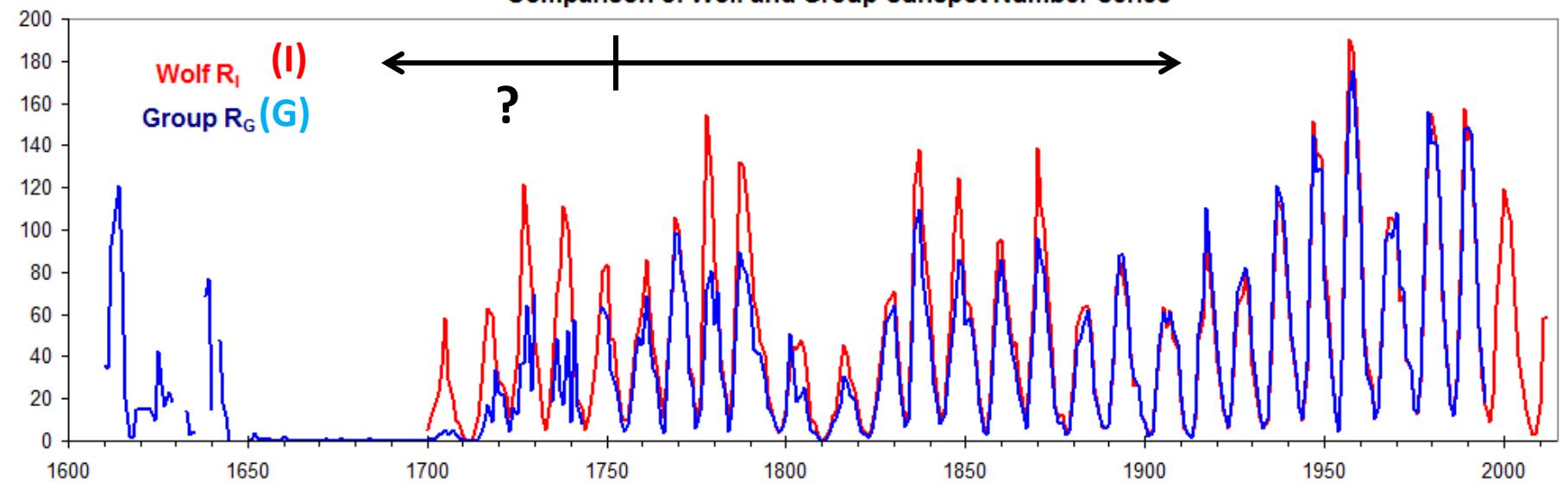
- 1) We think that we cannot use the backbone methodology in the early part of the sunspot number series.
- 2) Solar meridian observations should be used with extreme caution to evaluate past solar activity.
- 3) We have modified the HS98 data base, improving the quality of the compiled data (we hope!).
- 4) We will present a preliminary version of SSN series for the last four centuries, attaching the data from the revised database to the Staudach backbone.

# Attaching the Maunder Minimum to the Staudach Backbone

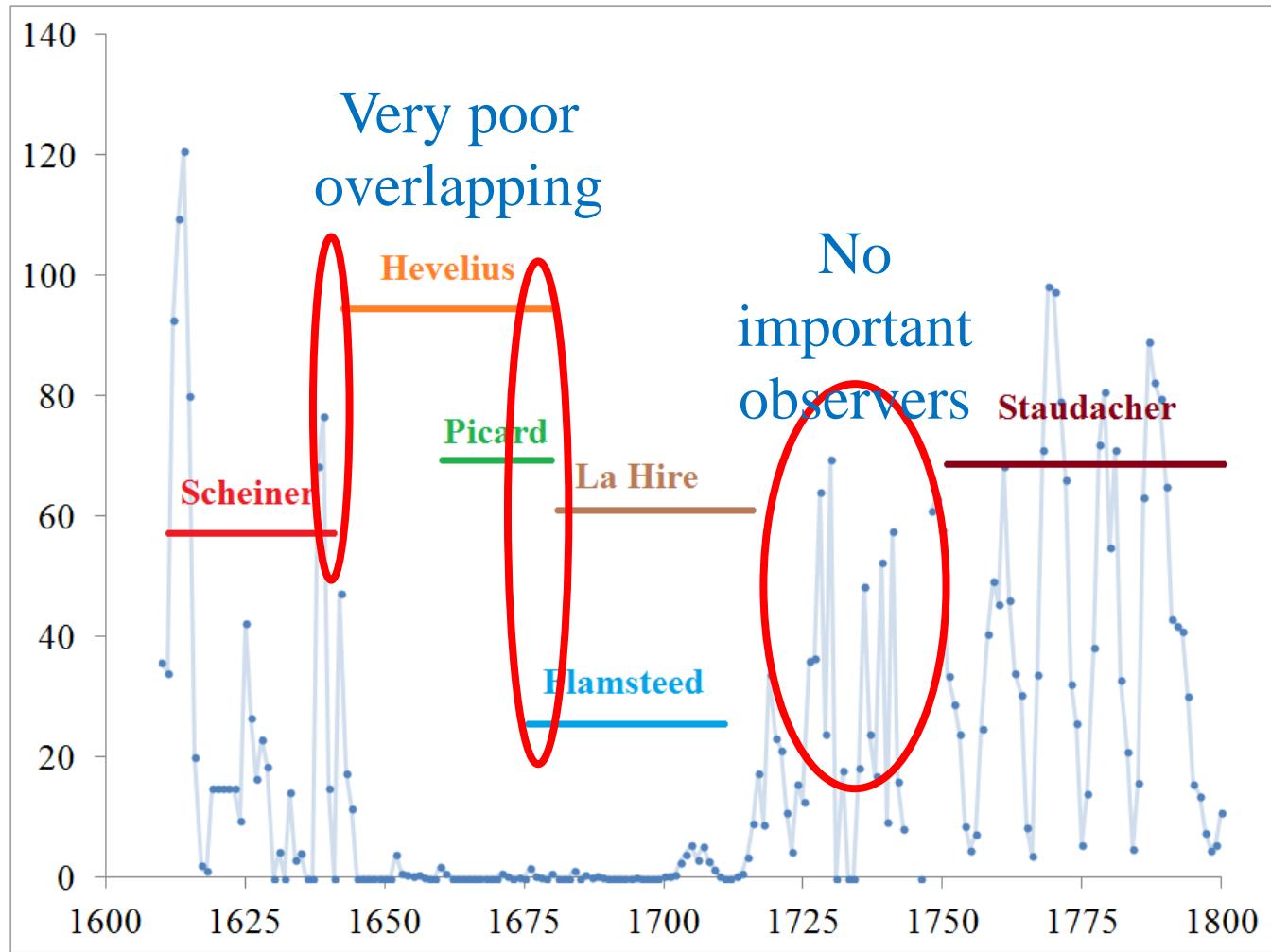


Clette et al. (2014)

Comparison of Wolf and Group Sunspot Number Series



# Maunder Minimum Vertebrae



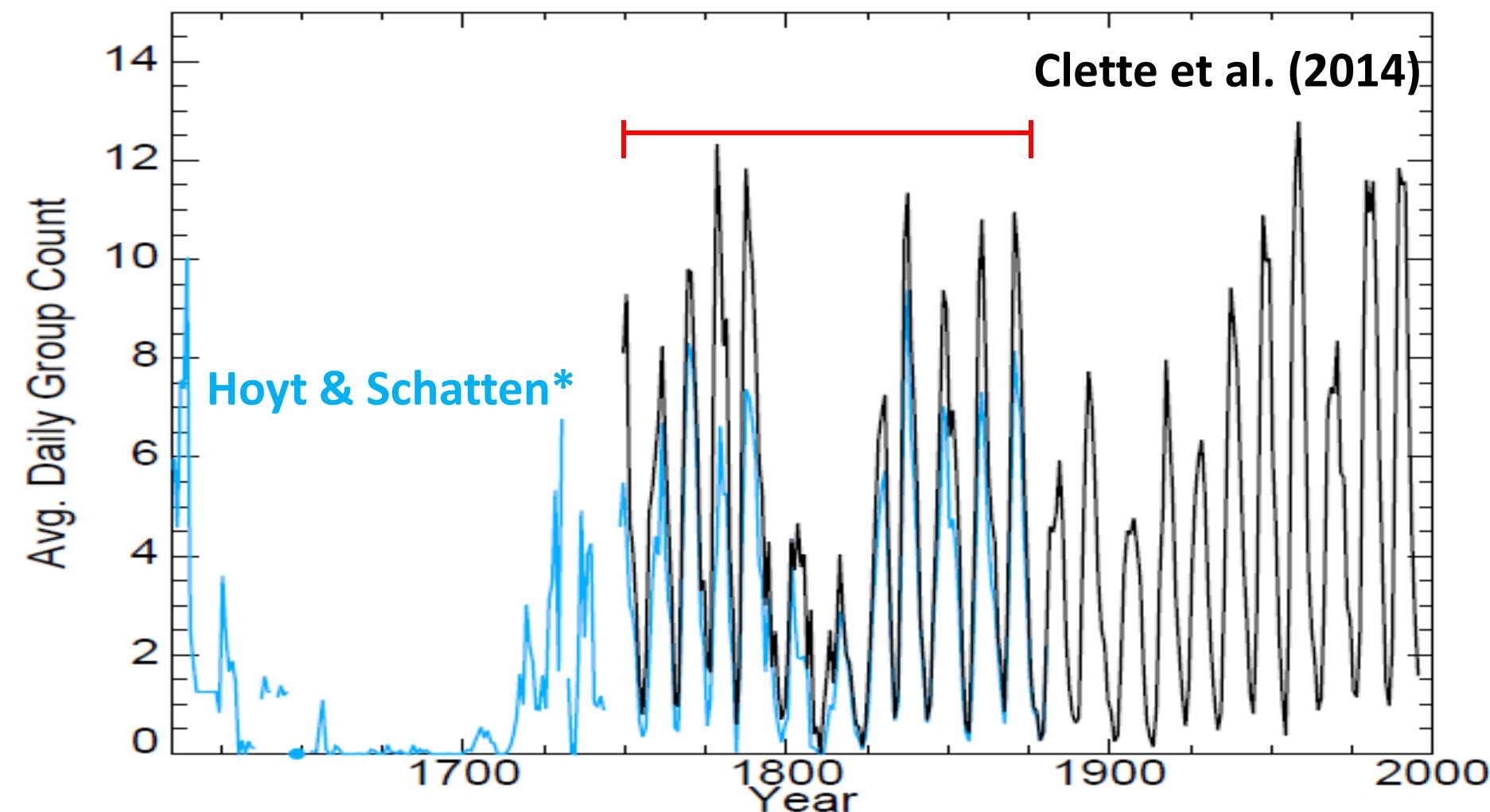
Year

HEVELIUS, J., DANZIG	
BOYLE, R., LONDON	7 1660-1660
CASSINI, G.D., PARIS	18 1671-1684
CASWELL (R.S.), LONDON	1 1684-1684
EIMMART, G.C., NUREMBERG	24 1677-1684
FLAMSTEED, J., CAMBRIDGE	45 1676-1684
FOGEL, M., HAMBURG	43 1661-1671
GULIELMINI, J.F., BONONIA	4 1684-1684
HOOK, R., LONDON	1 1676-1676
KIRCH, G., BERLIN	14 1684-1684
LA HIRE, PH., PARIS	13 1684-1684
LALANDE, MEMOIRES, PARIS	6 1672-1676
MACULA IN SOLE, 1678	1 1678-1678
MONTANARI, G., BONONIA	2 1676-1676
MOUTON, G., LYON	50 1659-1661
PETITUS, P., PARIS	9 1652-1676
PICARD, J., MONTPELLIER	2 1674-1674
<b>PICARD, J., PARIS</b>	<b>225 1660-1679</b>
PICARD/KEILL, PARIS	123 1653-1654
RICHER, CAYENNE	23 1672-1673
SIVERUS, H., HAMBURG	57 1671-1678
STETINI, LEIPZIG	4 1671-1671
UNKNOWN1/MAUNDER/JBAA	59 1652-1654
UNKNOWN2/MAUNDER/JBAA	2 1655-1655
WURZELBAUR, J.P., NUREMBERG	1 1684-1684

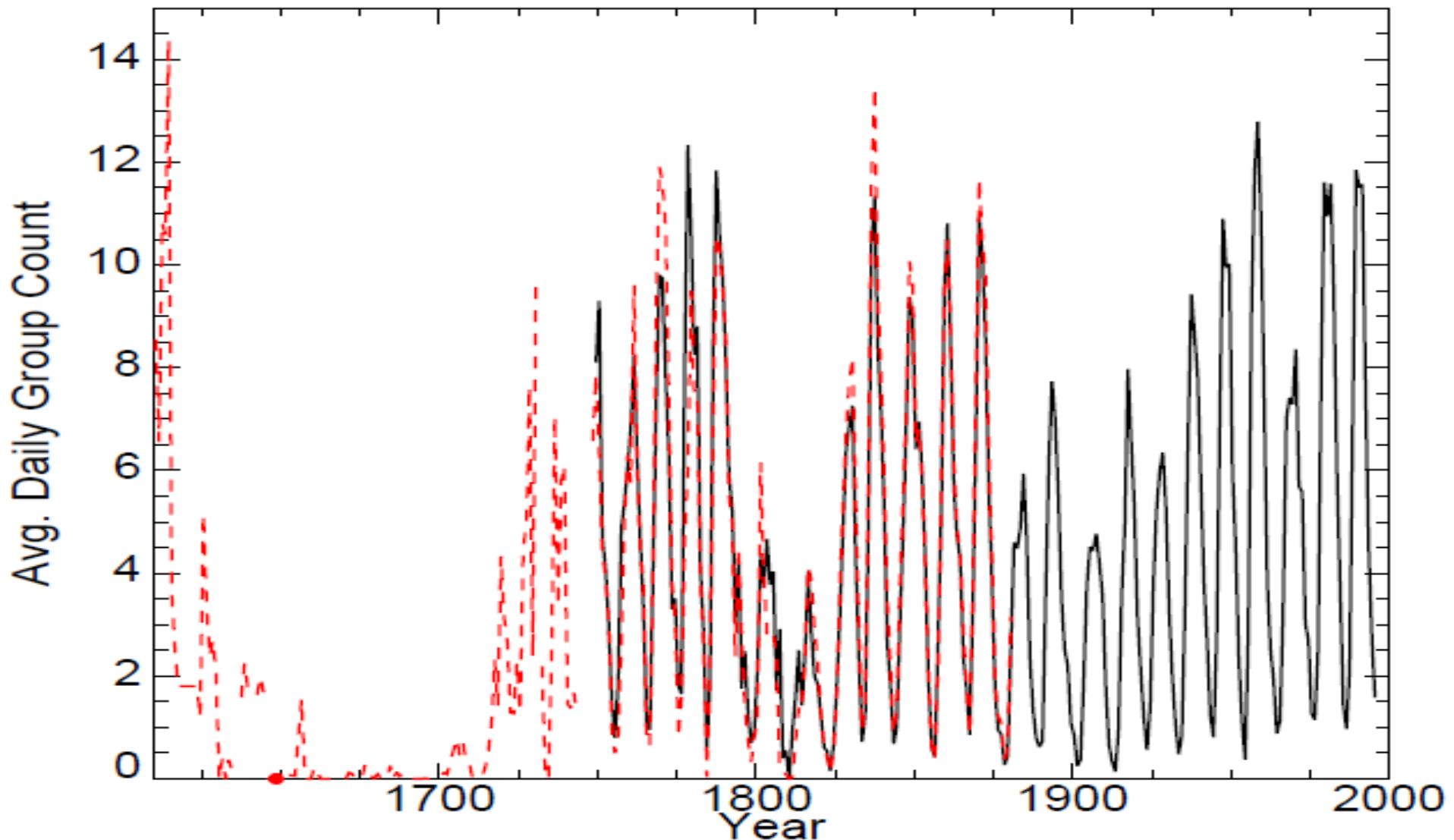
	No. of Sunspots	
	0	1
<u>Hevelius</u>	217	8
<u>Picard</u>	212	13

<u>Year</u>	<u>G.C.</u>	<u>Days</u>	<u>AD</u>	<u>NAD</u>
1645.5	1.8	1	1	0
1646.5	-1.0	0	0	0
1647.5	-1.0	0	0	0
1648.5	0.0	113	0	113
1649.5	-1.0	0	0	0
1650.5	-1.0	0	0	0
1651.5	-1.0	0	0	0
1652.5	0.1	58	2	56
1653.5	0.1	293	11	282
1654.5	0.1	331	7	324
1655.5	0.8	26	13	13
1656.5	1.6	15	13	2
1657.5	0.1	39	4	35
1658.5	0.0	23	0	23
1659.5	0.0	80	0	80
1660.5	0.2	294	47	247
1661.5	0.1	207	20	187
1662.5	0.0	51	0	51
1663.5	0.0	65	0	65
1664.5	0.0	50	0	50
1665.5	0.0	8	0	8
1666.5	0.0	93	0	93
1667.5	0.0	161	1	160
1668.5	0.0	144	0	144
1669.5	0.0	182	0	182
1670.5	0.0	136	0	136
1671.5	0.1	279	27	252
1672.5	0.1	199	11	188
1673.5	0.0	230	0	23

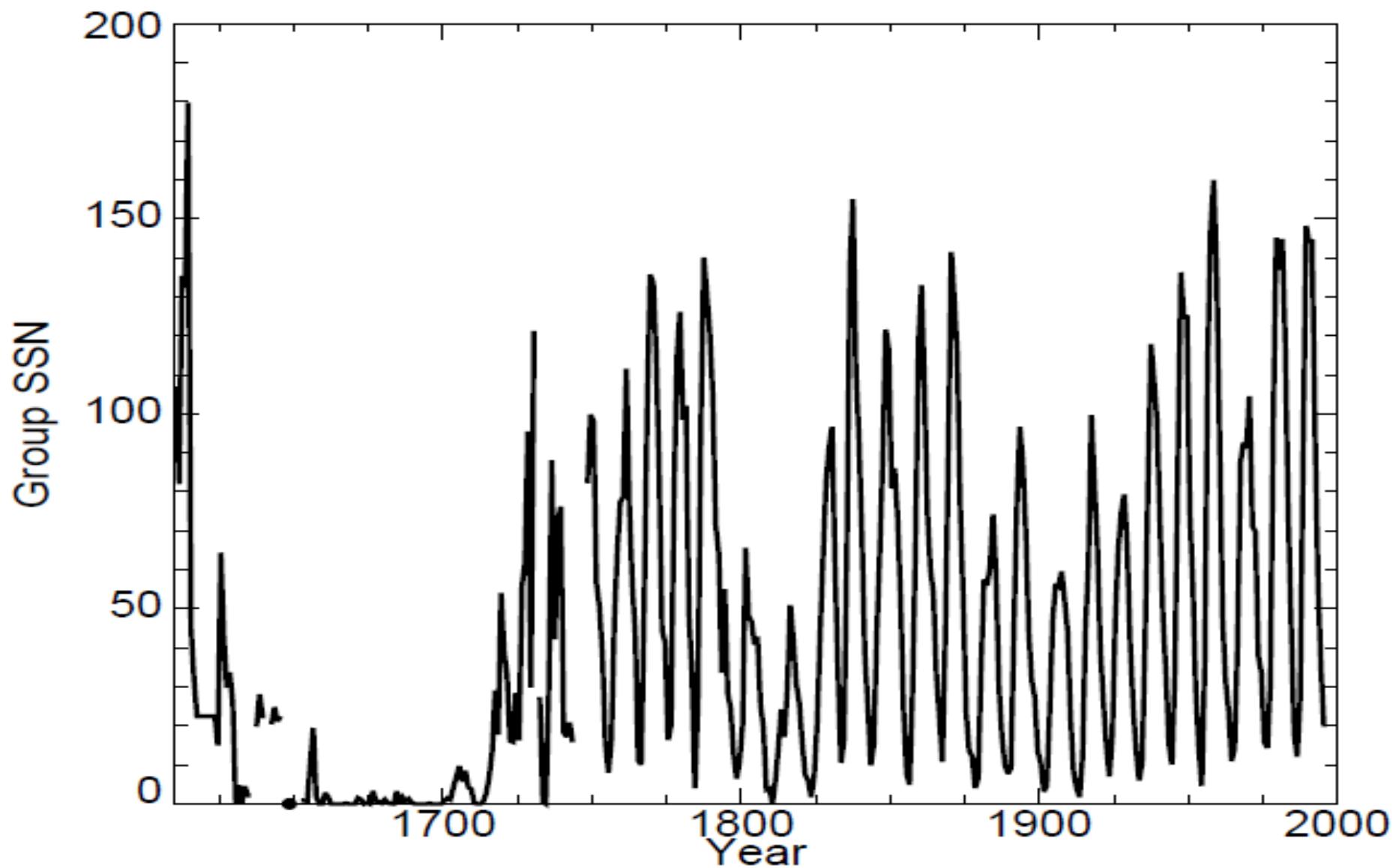
<u>Year</u>	<u>G.C.</u>	<u>Days</u>	<u>Obs.</u>	<u>AD</u>	<u>NAD</u>
1674.5	0.0	126		3	123
1675.5	0.0	193		0	193
1676.5	0.3	225		44	181
1677.5	0.1	197		10	187
1678.5	0.1	363		18	345
1679.5	0.0	167		2	165
1680.5	0.1	362		22	340
1681.5	0.0	360		7	353
1682.5	0.0	177		1	176
1683.5	0.0	222		0	222
1684.5	0.3	263		72	191
1685.5	0.0	200		0	200
1686.5	0.1	229		15	214
1687.5	0.1	222		9	213
1688.5	0.1	246		18	228
1689.5	0.0	257		5	252
1690.5	0.0	288		0	288
1691.5	0.0	282		0	282
1692.5	0.0	244		0	244
1693.5	0.0	263		0	263
1694.5	0.0	279		0	279
1695.5	0.0	335		4	331
1696.5	0.0	267		0	267
1697.5	0.0	266		0	266
1698.5	0.0	250		0	250
1699.5	0.0	284		0	284
1700.5	0.0	254		9	245



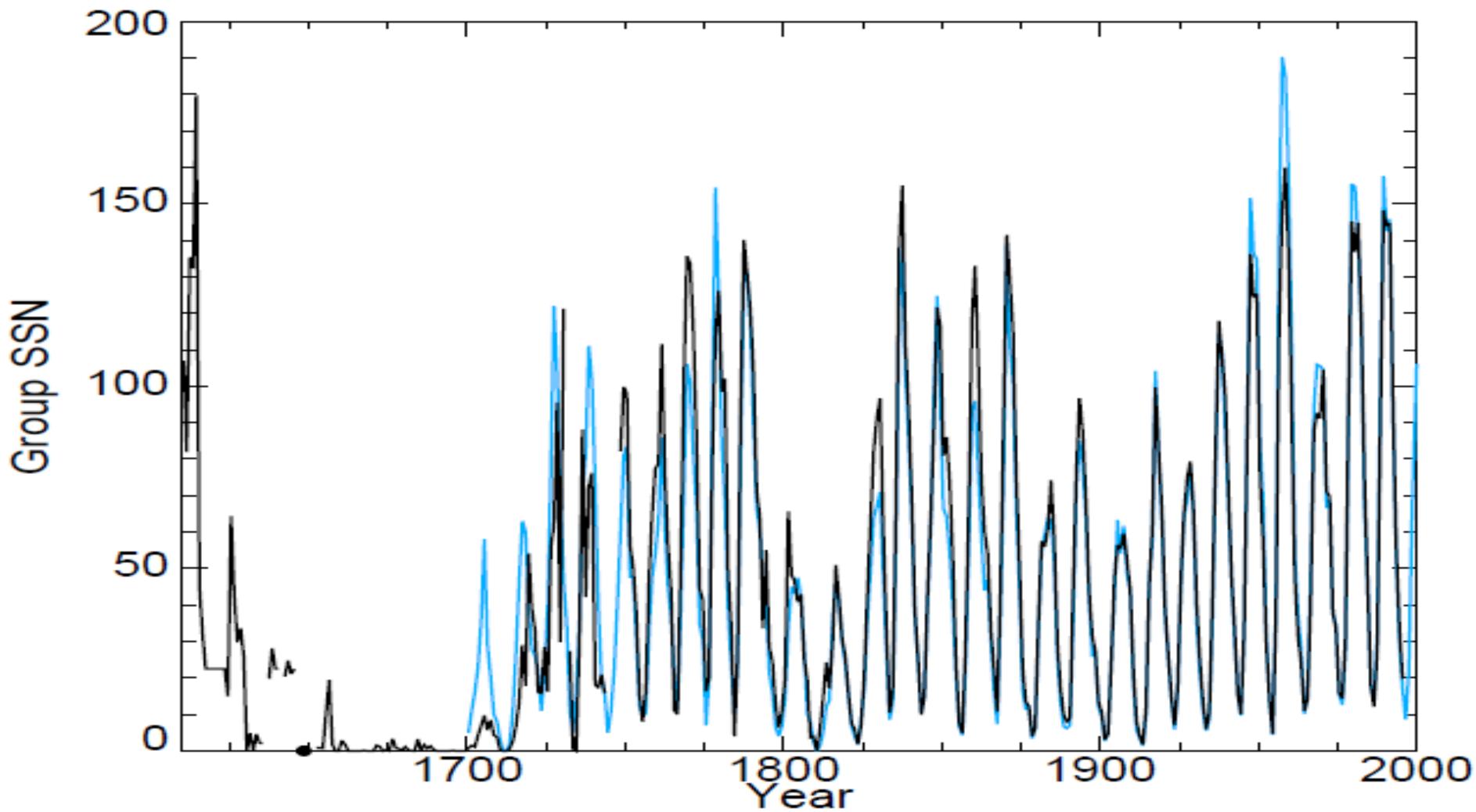
Comparison of H&S Group Counts and Clette et al. Group Counts  
For 1750-1880 overlap,  $\langle \text{Clette et al.} \rangle = 1.43 \times \text{H\&S}$



Average the two values for 1750-1880 &  
Multiply by 12.5 to convert to Wolf SSN



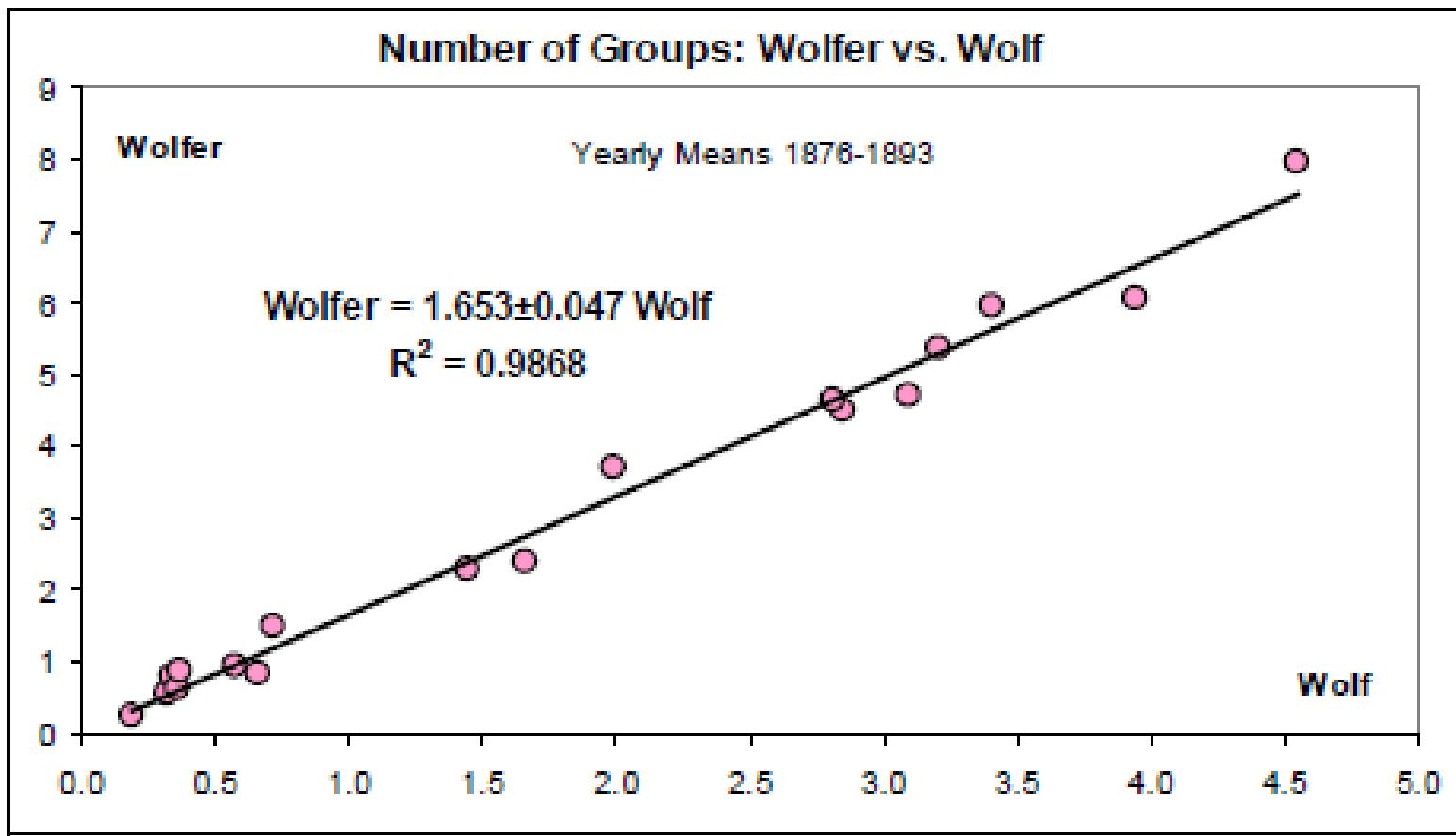
Preliminary Wolf SSN, 1610-Present



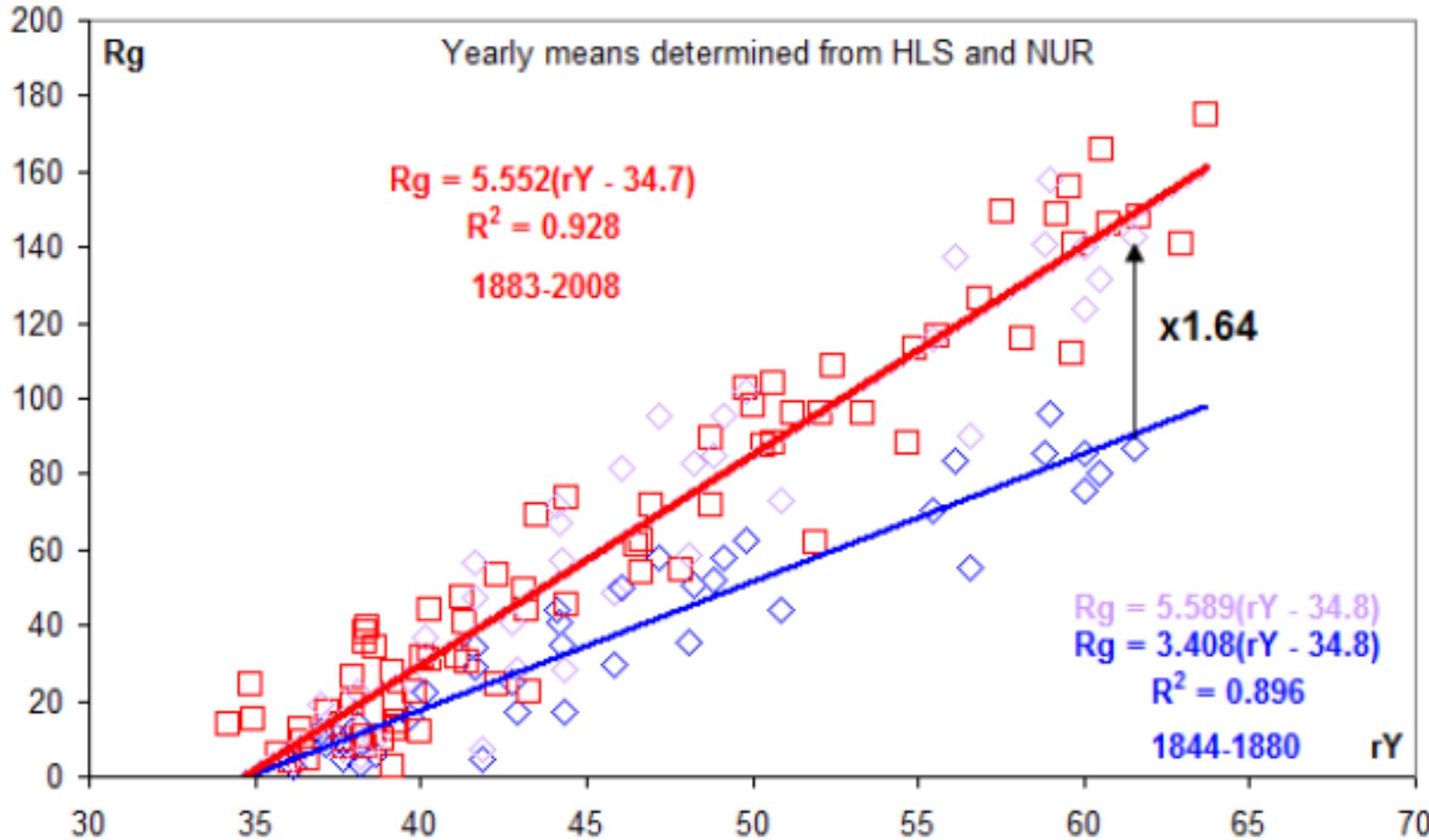
Comparison of Wolf Number with  $R_I$  (Brussels)

Justification for increasing the H&S Group SSN  
by ~50% before ~1880

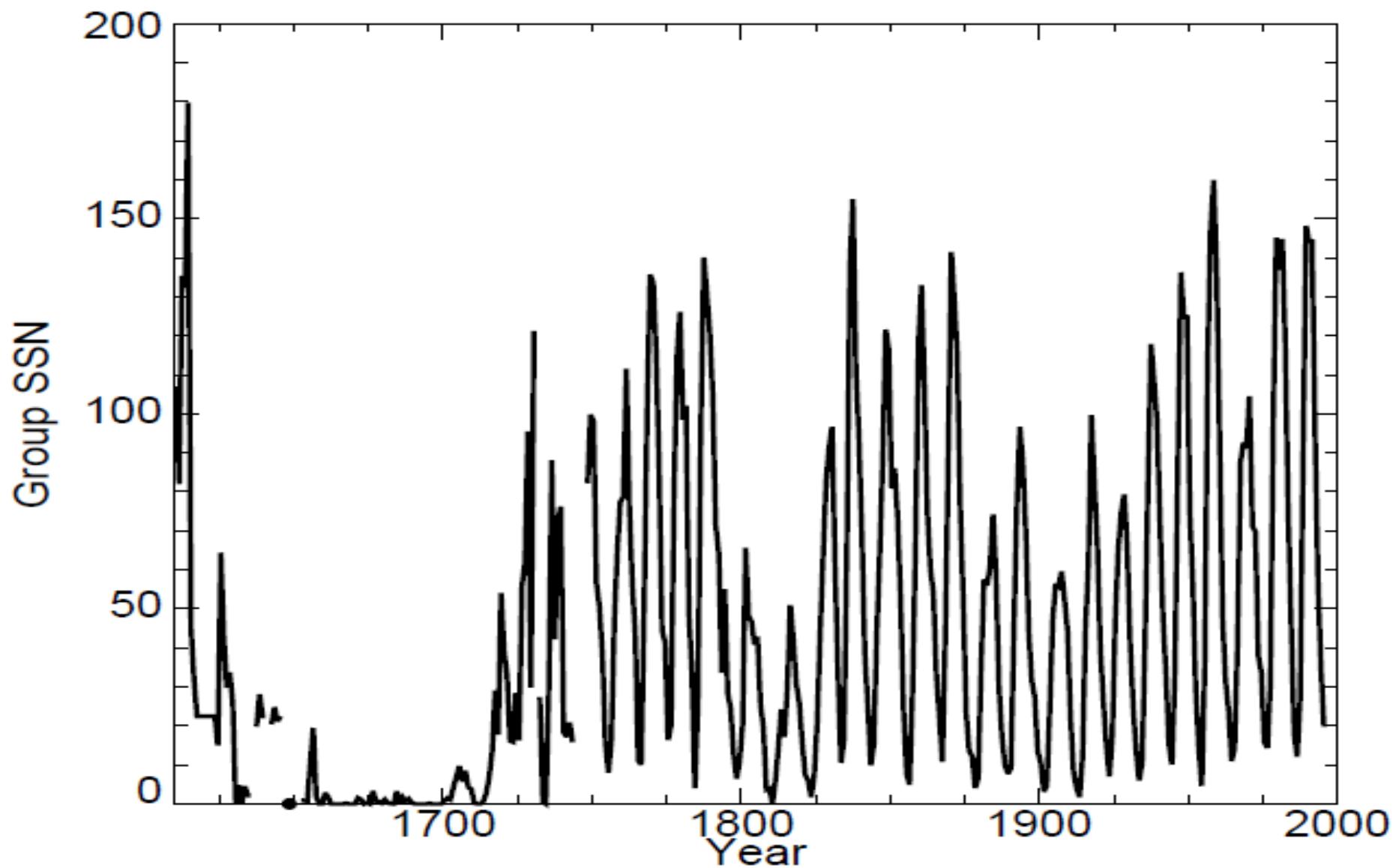
	K	$1\sigma$	Observer
298	<b>1.117</b>	<b>.090</b>	<b>WOLF, R.</b> ZURICH (1848-1893)
299	.937	.057	GREISBACH, T.J., ENGLAND
300	.758	.020	SESTINI, GEORGETOWN
...			
336	.796	.049	AGUILAR, MADRID
337	1.021	.063	MONTHLY WEATHER REVIEW
338	<b>1.094</b>	<b>.016</b>	<b>WOLFER, A.</b> ZURICH (1880-1928)



## Group Sunspot Number as a Function of Diurnal Range of East Component



The GSN need to be increased by a factor of 1.64 ( $\pm 0.15$ ) prior to  $\sim 1880$  unless the relationship between  $rY$  and GSN changed dramatically around that time.



Preliminary Wolf SSN, 1610-Present



